



## 21 Abstract

22 Methane seepage off the coast of Svalbard is demonstrated by active gas flares in the water  
23 column today and through precipitation of methane-derived authigenic carbonates (MDAC)  
24 over the past 160 000 years. Though submarine discharge of thermogenic methane is well  
25 documented, the geological history of past leakage is still debated, largely due to  
26 unconstrained free gas sources and seepage trigger mechanisms. We have assembled a  
27 high-resolution 3D petroleum systems model to evaluate the impact and charge of potential  
28 thermogenic hydrocarbons on the Vestnesa Ridge, NW Svalbard. We show that gaseous  
29 hydrocarbons, originating from Miocene age terrigenous organic matter, accumulates  
30 largely in ~2 million-year-old (Ma) sedimentary sequences underneath the pockmark system  
31 on the ridge's crest. These traps are constantly charged until present day. The supply of free  
32 gas to the gas hydrate stability zone (GHSZ) initiated the hydrate formation at ~3.0 Ma. We  
33 also show that gas leakage to the seafloor is governed by fault corruption of the  
34 hydrocarbon traps and not by excess pore fluid pressure. The onset of episodic seafloor  
35 seepage on Vestnesa Ridge can be associated with the first shelf edge glaciation of the  
36 Svalbard-Barents Sea ice sheet (SBIS), ~1.5 Ma ago. The results of the modelled petroleum  
37 system are consistent with the notion that repeated forebulge uplift and subsidence, due to  
38 cyclic SBIS build-up and decay, can be advocated as a mechanism that repeatedly caused  
39 extensional fracturing of the eastern Vestnesa Ridge segment. This fault damage affected  
40 modelled gas accumulations and led to the formation of hydrocarbon migration pathways to  
41 the seafloor. Repeated tapping into hydrocarbon reservoirs due to (a) fracture formation  
42 promoted by glacial isostatic adjustments (GIA) and (b) fracture re-activation could explain

43 recent observations of multiple seepage events on Vestnesa Ridge during episodes of  
44 intense cooling over the past 160.000 years.

## 45 1 Introduction

46 The dynamics of Arctic methane seepage in the eastern Fram Strait has been linked to gas  
47 hydrate dissociation regulated by several mechanisms: long term and seasonal climatic  
48 changes (Westbrook et al., 2008; Berndt et al., 2014), retreat of the Svalbard-Barents Sea ice  
49 sheet (SBIS) (Portnov et al., 2016) and isostatic adjustment after the Last Glacial Maximum  
50 (LGM) (Schneider et al. 2018, Wallmann et al., 2018, Ferré et al., 2020). However, gas  
51 hydrates are not widespread in areas of abundant seepage near the shelf break; hence gas  
52 hydrate dissociation as sole explanation for seepage in the region remains elusive (Smith et  
53 al. 2014, Pape et al. 2019). Instead, a link between seepage sites, gas chimneys and near-  
54 surface faults from seismic data along the Vestnesa Ridge point towards a close relation  
55 between episodic seepage and fault dynamics (Plaza-Faverola et al., 2015).

56 While present day seepage distribution along the Vestnesa Ridge may be controlled by fault  
57 dilation under a tensile tectonic stress regime due to spreading at the Molloy and Knipovich  
58 ridges (Plaza-Faverola and Keiding, 2019), historical seepage periodicity through faults was  
59 possibly influenced by glacio-isostatic adjustments (GIA) of the SBIS (Schneider et al., 2018;  
60 Plaza-Faverola et al., 2019; Himmler et al., 2019). It has been hypothesized that since  
61 intensification of the Northern Hemisphere glaciation, ~2.7 million years ago, episodic  
62 release of Arctic methane at the seabed has increased (Plaza-Faverola et al. 2015); a fact  
63 that is corroborated by recent fluid migration modelling results in the Arctic-Atlantic  
64 gateway (Knies et al. 2018). On younger time scale, Schneider et al. (2018) recently  
65 summarized episodic seepage events in the gateway region during build-up and decay of the

66 SBIS during the LGM implying a close relationship between GIA and fault-reactivation to  
67 open, highly permeable conduits and enhanced methane flux. Recently, Himmler et al.  
68 (2019) suggested that frequent methane release events in the region have occurred since  
69 the Penultimate Glacial Maximum (PGM), about 0.160 Ma ago. Also, these events correlate  
70 to ice sheet dynamics of the SBIS. The geochronology of these events indicates that some  
71 episodes of methane release coincide with dramatic climatic shifts in the Northern  
72 Hemisphere expressed by so-called Heinrich Events (Hemming, 2004) – thick layers of ice-  
73 rafted debris deposited in the North Atlantic during glacial times because of massive  
74 discharges of icebergs from the Laurentide ice sheet – implying a causal relationship  
75 between large-scale ice sheet dynamics and methane release in the Arctic.

76 Although large uncertainties exist on the extent and volume of circum-Arctic ice sheet  
77 beyond the last glacial period (Patton et al. 2015, 2016), it is hypothesized here, that the  
78 source of seeping methane off western Svalbard is of pre-glacial origin (Miocene) and the  
79 transfer of methane into the seepage system may have been triggered and facilitated by  
80 various mechanisms during the Plio-Pleistocene. These mechanisms comprise (1) excess  
81 fluid pressure build-up, (2) hydrocarbon migration along large scale fault surfaces, (3)  
82 persistent upward hydrocarbon migration across the stratigraphic inventory, (4)  
83 hydrocarbon leakage loss from shallow gas accumulations or (5) various combinations of  
84 them. We are testing our hypothesis by applying an innovative 3D hydrocarbon migration  
85 modelling approach on one of the most active methane seep sites in the Arctic, the  
86 Vestnesa Ridge, NW Svalbard.

## 87 2 Geological setting

88 The Vestnesa Ridge is a ~100 km long, boomerang-shaped and pockmark covered  
89 submarine feature offshore western Svalbard (Fig. 1). It hosts sediments of >5 km in  
90 thickness and rest on top of hot (heat flow  $>115\text{mW/m}^2$ ) and relatively young oceanic crust  
91 (<20 Ma), near the ultraslow spreading Molloy and the Knipovich ridges in the Fram Strait  
92 (Eiken and Hinz, 1993; Vogt et al., 1994; Hustoft et al., 2009 and Bünz et al., 2012). Slow  
93 spreading of the Molloy and Knipovich ridges led to the onset of the Fram Strait opening  
94 about 33 Ma ago (Engen et al., 2008). Deep marine conditions and water mass circulations  
95 were established in the Fram Strait in the Miocene, ca. 17 – 10 Ma ago (Jakobsson et al.,  
96 2007; Ehlers and Jokat, 2009), and shaped the environmental conditions for the evolution of  
97 bottom current-driven sedimentary drifts (Eiken and Hinz, 1993; Johnson et al., 2015).

98 Three main stratigraphic units (YP1-YP3) are designated according to their correlation with  
99 Ocean Drilling Program (ODP) sites at the Yermak Plateau (Eiken and Hinz, 1993,  
100 Mattingsdal et al. 2014). Miocene – Pliocene age syn and post-rift deposits overlying the  
101 oceanic crust (YP1) are superseded by contourites (YP2). By the onset of the Pleistocene  
102 around 2.7 Ma the contourites became intermixed with glaciogenic debris flow deposits  
103 sourced from the East (YP3) (Howe et al., 2008).

104 Restricted to the youngest stratigraphic sequence (YP3), a gas hydrate system underlain by a  
105 free-gas zone occurs along-strike the entire Vestnesa Ridge. This system is permeated by a  
106 series of fault bound gas chimneys which terminate in pockmarks at the seafloor (Hustoft et  
107 al., 2009; Petersen et al., 2010; Plaza-Faverola et al., 2015, Panieri et al., 2017 and Plaza-  
108 Faverola and Keiding, 2019). Active methane discharge from pockmarks into the water  
109 column is recorded on the eastern ridge segment today (Bünz et al., 2012; Smith et al.,

110 2014). Multiple proxies such as radiogenic and stable isotopes, foraminifera and methane-  
111 derived authigenic carbonate (MDAC) crusts in the geological record indicate episodic  
112 seepage over the past 160 ka (e.g. Ambrose et al., 2015, Panieri et al., 2017, Szttybor and  
113 Rasmussen, 2017, Schneider et al., 2018, Himmler et al., 2019 ). However, seismic data  
114 showing seepage features such as gas chimneys and buried pockmarks seem to indicate that  
115 methane release has been prevailed since the intensification of the Northern Hemisphere  
116 glaciation, about 2.7 million years ago (Plaza-Faverola et al., 2015; 2019). Contrasting, no  
117 seepage activity is observed on the western ridge today though foraminiferal records  
118 indicate methane seepage activities about 14 – 13 and 10 ka ago (Consolaro et al., 2014).  
119 Isotopic  $\delta^2\text{H}$  and  $\delta^{13}\text{C}$  signatures in methane and ethane from dissolved gas and hydrates  
120 samples (Faverola et al., 2017, Pape et al., 2019, Sauer et al., 2021) collected at the Lunde  
121 and Lomvi seep sites (Fig. 1) evidence a deep thermogenic source of the hydrocarbons  
122 charging the hydrate and seep system from below the gas hydrate stability zone (GHSZ) at  
123 the eastern ridge segment.

## 124 3 Material and Methods

### 125 3.1 Petroleum systems model

126 Petroleum systems modelling (PSM) is a well-established method in hydrocarbon  
127 exploration and numerically reproduced processes of hydrocarbon generation, migration,  
128 trapping and leakage within sedimentary basins over geological times. We employ the Migri  
129 simulator (<http://www.migris.no/technology>), a fast commercial PSM software. Modelled  
130 properties are treated dynamically, meaning their values change through geologic time  
131 unless stated otherwise. Details on the methods for modelling the individual processes are  
132 described in [Sylta \(2004\) and Sylta \(2005\)](#). The vertical pore entry pressure distribution

133 within the layers is evaluated to determine whether primary migration out of the source  
134 rocks should be directed upward or downward. Over time, entry pressure values in the grid  
135 elements are calculated from the permeabilities, which are derived from the lithological  
136 composition and burial depths. Secondary migration is modelled as mostly buoyancy-driven  
137 within permeable systems underneath flow barriers where the direction of flow is governed  
138 by the steepest gradient of ascent. Lateral entry pressure differences within the flow unit  
139 may, however, cause the flow to be deflected. Lateral migration may also be stopped or  
140 deflected when the flow encounters a shale pinch-out barrier with higher entry pressures  
141 than more porous strata. Migrating hydrocarbons are modelled to accumulate in traps. Gas  
142 and subsequently oil are modelled to leak out of the trap according to the entry pressure  
143 distributions modelled in the sealing rocks (e.g. in shales). Both, the trapped hydrocarbon  
144 column heights, and the seal permeabilities determine the rates and areas of leakage from  
145 traps.

146 The 3D petroleum systems model and associated initial boundary conditions are of the same  
147 derivation as previously used to model the methane seepage along the 2D seismic line  
148 across the Vestnesa Ridge (Knies et al. 2018) where a more complete description of the  
149 model setup and implementation can be found. Details on the key changes to expand the  
150 2D into a 3D-PSA model (Fig. 1) are given in the enclosed supplements. Key petrophysical  
151 input parameters are given in Tab. 1. The main elements of this PSM comprise the basin  
152 configuration and stratigraphy, burial and thermal history models as well as models for  
153 lithological and organic matter inventory in the source rocks and are briefly summarized  
154 here.

155 3.1.1 Basin configuration and burial history model

156 Four main depth horizons and the seafloor were interpreted from several 2D seismic lines  
157 (CAGE-VR2D-005, CAGE-VR2D-006, SV 3-97 , Svalex P11, Svalex P12) and one 3D P-cable  
158 seismic volume (Plaza-Faverola et al., 2015 and references therein, Knies et al., 2018). These  
159 include: Top basement (assigned age of 300 Ma to avoid impact on the basin modelling), the  
160 6.2 Ma regional biostratigraphic datum for the Nordic Seas and Arctic Ocean (Matthiessen  
161 et al., 2009) and the chronostratigraphic framework established for the Yermak  
162 Plateau/western Svalbard including the Matuyama/Gauss chron boundary (~2.6 Ma) and  
163 the ~1.5 Ma old regional seismic reflector (Knies et al., 2009; Mattingsdal et al., 2014).

164 Details on the depth conversion are given in the enclosed supplements. Additional depth  
165 horizons were constructed using lithology-age-depth models of ODP Hole 909C (Winkler et  
166 al., 2002; Wolf-Welling et al., 1996) and Hole 986D (Jansen et al., 1996) with assigned ages  
167 to the model of 0.02 Ma, 0.043 Ma, 0.148 Ma, 2.1 Ma, 3.2 Ma, 4.39 Ma, 12.05 Ma, 15.18  
168 Ma, 17.37 Ma, and 18.5 Ma (Fig. 2). Burial history reconstruction was performed using a  
169 back-stripping approach (Allen and Allen, 2005) with paleo-bathymetries inferred to follow  
170 present-day marine trends, assuming constant tectonically induced subsidence. Thicknesses  
171 and depth of units over time are derived from decompaction modelling using the concept of  
172 Sclater and Christie (1980). The permeabilities of the units are calculated dynamically over  
173 time and range from impermeable to 5000 mD. They are based on the lithologies and  
174 porosity estimates from the decompaction calculation (Tab. 2). The modelling is performed  
175 over 10 timesteps between 12.1 Ma and present-day (Tab. 2).

### 176 3.1.2 Thermal history model

177 The thermal history model infers the highest geothermal gradients at the Molloy Transform  
178 Fault (MTF) location and decreasing gradients away from the MTF in SW and NE directions  
179 (Vanneste et al., 2005). Data from Crane et al. (1992) and ODP sites 908, 909, 910 and 912  
180 (Stein et al., 1995) were jointly used to interpolate geothermal gradient maps across the  
181 area of interest (AOI). Geothermal gradient values range from 55 °C/km to maximum  
182 120 °C/km at the MTF (Tab. 1 and Fig. S1). Seafloor temperatures are set uniformly to -1°C  
183 (Plaza-Faverola et al. 2017). Inferred geothermal gradients agree with heat flow values  
184 reported by Hustoft et al. (2009) and Smith et al. (2014) west of Svalbard, considering  
185 thermal conductivities of silt and shale lithologies. High sedimentation rates during  
186 Pliocene/Pleistocene times at the Vestnesa Ridge exerted a cooling effect in the upper  
187 sedimentary layers (Plaza-Faverola et al., 2017). We note that the resultant transient  
188 changes in the thermal regime challenge the assumption of a constant geothermal gradient  
189 over time and, hence, impact the correct approximation of the timing of source rock  
190 maturation. To account for this, we apply the equation of Palumbo et al. (1999) to the  
191 gradient model and thereby better assess the timing of the source rock maturation due to  
192 rapid sedimentation. Modelled burial and thermal histories for four locations across the AOI  
193 (including vitrinite reflection (VR) calibration data at ODP Site 909) are shown in the  
194 enclosed supplements (Figs S2 and S3).

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197 3.1.3 Lithological model

198 Sand-shale-carbonate ratios observed in ODP Hole 909C (Wolf-Welling et al., 1996) and Hole  
199 986D (Butt et al., 2000; Knies et al., 2009) are combined into a composite lithology log  
200 whereby age and lithology information from ODP Hole 986D are projected onto the model's  
201 age-depth-lithology relation at the Vestnesa Ridge (Fig. 3). We expect that Pleistocene,  
202 glacially-derived sediments there are more coarse-grained compared to pre-glacial  
203 sediments (with age >2.7 Ma) and are comparable with the lithological composition of ODP  
204 Hole 986D (Butt et al., 2000). Consequently, lithostratigraphic information from ODP Hole  
205 909C (for stratigraphic ages from 17.4 Ma to 2.6 Ma) and the projected ODP Hole 986D (for  
206 stratigraphic ages from 2.6 Ma to 0 Ma) were combined to populate the model with a  
207 laterally invariant sand-shale-carbonate ratio stratification, between the top basement and  
208 the seafloor reflectors. The entire lithological model is stratified according to borehole  
209 information and therefore accounts for vertically changing proportions of shale, sand, and  
210 carbonate (Fig. 3) Model porosities and permeabilities are dynamically derived from the  
211 lithology properties in the cells (Tab. 1) meaning that units with low shale and high sand  
212 contents are characterized by higher porosities and permeabilities compared to high shale  
213 content strata at a similar burial depth.

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219 Table 1: Petrophysical input parameters to the petroleum systems model

Parameter/variable	Values/units/dependencies	Physical meaning	Reference
Spatial resolution	50 m x 50 m avg. 10 – 30 m, depends on $N_{sublayer}$ .	Area represented by a grid cell in a 3D element. Height of a 3D element	
$N_{horizon}$	15 in total, 5 based on seismic interpretation, 10 constructed from seismic horizon- and well-data (ODP holes 909C and 986D)	Number of depths horizons employed in the layer model	
$N_{sublayer}$	2 – 20 per layer, depending on lithology variation and layer thickness	Number of discrete lithological subunits in geo-model layer between 2 depth horizons	
$N_{time}$	10 between 12.1 Ma and 0 Ma	Number of model time steps	
$dT/dz$	55 - 120 °C/km, laterally variable	Geothermal gradient, map in Fig. S1	Crane et al., 1992 Stein et al., 1995 Hustoft et al., 2009, Johnson et al., 2015, Smith et al., 2014,
$V_{sh}, V_{ca}, V_{ss}$	0 – 1, fraction	Fraction of shale, carbonate, and sand with $V_{sh}+V_{ca}+V_{ss} = 1$	
$\phi$	0.0077 – 0.5780, fraction, depending on lithology and burial depth	Porosity of the modelled lithology	
$K$	$10^{-7} - 5 \cdot 10^3$ mD, depending on lithology and burial depth, cells in sealing gas hydrate zone = $1 \cdot 10^{-5}$ mD	Permeability of modelled lithology	
$P_c$	$2 \cdot 10^{-2} - 2.9 \cdot 10^2$ MPa, depending on hydrocarbon phase and $K$ , cells in sealing gas hydrate zone = $2.1 \cdot 10^1$ MPa	Capillary entry pressure of pore space	

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221 3.1.4 Source rock model

222 Two organically enriched and thermally immature siliciclastic intervals of middle Miocene  
 223 age are encountered in ODP Hole 909C (Myhre et al., 1995). The organic matter is  
 224 terrigenous nature and originated from the bordering basin flanks of AOI<sub>west</sub> (Hovgaard  
 225 Ridge) and AOI<sub>east</sub> (Svalbard Platform) (Knies and Mann, 2002). Two source rock units with a  
 226 type III kerogen were defined between the depth horizon intervals 17.37 Ma and 15.18 Ma  
 227 (source rock unit 1) and 15.18 Ma and 12.05 Ma (source rock unit 2) from direct observation  
 228 in ODP Hole 909C (Knies and Mann, 2002, Stein et al., 1995). Source rock unit 1 employs  
 229 laterally variable total organic carbon (TOC) and hydrogen index (HI) values both in the  
 230 AOI<sub>east</sub> and AOI<sub>west</sub> (Fig. 1) in accordance with the organic facies distribution model of Knies  
 231 and Mann (2002). Employed TOC values range from 0 to 3.5 wt% and HI values span

232 between 20 and 200 mgHC/gTOC with maximum effective thicknesses being limited to 75 m  
 233 (Tab. 2). Further details on the model derivation and lateral value distributions are given in  
 234 the enclosed supplements (Fig. S4). For the younger source rock unit 2, uniform values for  
 235 TOC of 1.0 wt% and HI of 50 mgHC/gTOC are used, as observed in ODP Hole 909C (Myhre et  
 236 al., 1995). Its effective thickness is delimited to a maximum of 100 m (Tab. 2).

237 Table 2: Source rock input parameters to the petroleum systems model

Parameter/variable	Values/units/type	Physical meaning	References
<i>Source rock unit 1 – age: 17.37 – 15.18 Ma</i>			
TOC <sub>0</sub> [1]	0-3.5 wt%	Total organic carbon in the source 1 prior to thermal alteration	Knies and Mann, 2002
HI <sub>0</sub> [1]	0-200 mgHC/gTOC	Hydrogen index of kerogen in source 1 prior to thermal alteration	Stein et al., 1995
Thickness [1]	up to 75 m	Effective thickness of source 1	
<i>Source rock unit 2 – age: 15.18 – 12.05 Ma</i>			
TOC <sub>0</sub> [2]	1 wt%	Total organic carbon in the source 2 prior to thermal alteration	Stein et al., 1995
HI <sub>0</sub> [2]	50mg/gTOC	Hydrogen index of kerogen in source 2 prior to thermal alteration	
Thickness [2]	up to 100 m	Effective thickness of source 2	
<i>Source rock characterization</i>			
Kerogen type [all]	Terrigenous, type III	Kerogen type classification of the organic matter in source rocks [1, 2]	Stein et al., 1995
Kinetic scheme [all]	Behar type III	Kinetic scheme for the kerogen to hydrocarbon transformation for all source rocks [1, 2]	Behar et al., 1997

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### 239 3.1.5 Methane hydrate stability zone model

240 A model (for numerical details see supplements) to determine the depth of the base of the  
 241 GHSZ and incorporate the GHSZ as a sealing barrier to migrating hydrocarbons over geological  
 242 times has been implemented. The concept employs the CSMHyd program (Sloan and Koh,  
 243 2007) for calculating the P-T phase boundary curves for gas hydrates with mixed source gas  
 244 compositions. We employed two end members of source gas composition from Plaza-  
 245 Faverola et al. 2017 (Scenario A: 100% C<sub>1</sub> and Scenario B: 95.6% C<sub>1</sub>, 2.06% C<sub>2</sub> and 2.32% C<sub>3</sub>-  
 246 C<sub>6</sub> from sample GC19, see Table S1 in supplements) to test the potential impact of variable  
 247 GHSZ thicknesses on hydrocarbon migration patterns over time. Specifically, it was tested if  
 248 the directions of lateral HC migration routes at shallow depth will be deflected by a GHSZ

249 barrier and if such barrier is suited to accumulate continuous columns of gaseous  
250 hydrocarbons beneath it.

### 251 3.1.6 Faults and fault leakage model

252 Two types of fault sets were interpreted in the seismic dataset (Fig. 4). The first set (set 1)  
253 consists of fault outline polygons collated from fault sticks interpreted within 2D seismic  
254 lines and corresponding fault scarps exposed at seabed surface. Faults of this set have a  
255 regional extent in the AOI, occur on the southern flank of the Vestnesa Ridge towards the  
256 northward end of the Knipovich Ridge (Fig. 1) and trend N-S to NW-SE. They are steeply  
257 dipping at angles between 50 to 75 degrees between the seafloor and the 3.2 Ma depth  
258 horizon (Fig. 4). Except for their surface fault scarp, they do not exhibit any distinct vertical  
259 offset at horizons below the seafloor at the resolution of the available seismic surveys.  
260 Timing of the formation for these faults is unknown. However, based on the relative age of  
261 the horizons truncated by these faults, we assume that faulting and re-activation have  
262 occurred between  $\sim 3$  Ma at the earliest and today evidenced by the seabed fault scarps.  
263 Fault set 2 is restricted to the crestral zone of the Vestnesa Ridge and is partially imaged by  
264 the 3D P-cable dataset (Fig. 1). These faults have a very local occurrence and strike NNW-  
265 SSE to NW-SE, strongly paralleling the trend of the south-eastern Vestnesa Ridge segment.  
266 They have nearly vertical surfaces with dip angles  $>80$  degrees and extend from the seabed  
267 to the YP3/YP2 boundary ( $\sim 2.7$  Ma) (Plaza-Faverola et al., 2015). According to Plaza-Faverola  
268 and Keiding (2019), they appear to be extensional in nature and have a linkage to seismic  
269 gas chimney expressions in the subsurface. Their formation age is unknown but relative  
270 relations to fractured horizons suggest an origin age between less than  $\sim 2.7$  Ma age and  
271 today (Plaza-Faverola et al., 2015).

272 In the hydrocarbon migration model, fault set 1 was defined to permit up-dip hydrocarbon  
273 migration (fault leakage) from deeper into shallower carrier horizons. The fault leakage  
274 concept permits a pre-defined fraction of hydrocarbons (volume) from a migration stringer  
275 passing along a fault surface to move upward and exiting the fault at a shallower level if  
276 suited pore entry pressure and permeabilities for lateral hydrocarbon (HC) migration are  
277 met. If no favorable conditions are encountered at shallower levels, no fault leakage is  
278 modelled, and all HC volumes remain in the original migration stringers at depth.

279 Fault set 2 is too local with respect to the grid resolution to be included in the fault leakage  
280 model and hence no leakage model was applied .

### 281 3.2 Local stress considerations

282 A local stress analysis was performed to evaluate the stress impact of modelled gas  
283 accumulations on the overburden stability and their relation to documented methane  
284 seepage history at the western Vestnesa Ridge crest. To evaluate the local stresses in a  
285 passive margin setting, we assume that  $\sigma_1 = \sigma_v$  (at or near vertical), and the minimum  
286 principle stress,  $\sigma_3 = \sigma_h$  (at or near horizontal). We calculate  $\sigma_v$  (total vertical stress) from wet  
287 bulk density data of ODP Hole 986C and Hole 986D, and use the poro-elastic formulation of  
288 Engelder and Fischer (1994) to estimate the total horizontal stress,  $\sigma_h$ , considering the  
289 Poisson's effect and the pore pressure increase as to:

290

$$291 \quad \sigma_v = \int_0^z \rho_{bulk} g dz \quad (3)$$

292

$$293 \quad \sigma_h = \left[ \frac{\nu}{1-\nu} \right] \sigma_v + \alpha \left[ \frac{1-2\nu}{1-\nu} \right] (P_{water} + P_{gas}) \quad (4)$$

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295 where  $z$  is the overburden thickness above a hydrocarbon trap or at the BSR,  $\rho_{bulk}$  is the  
296 sediment bulk density,  $g$  is gravitational acceleration,  $\nu$  is the Poisson's ratio of sediment at  
297 the top of a trap or at the BSR,  $\alpha$  is the Biot-Willi poroelastic coupling constant of the  
298 effective stress, and  $P_{water}$  and  $P_{gas}$  are the fluid pressures of the water and gas phases,  
299 respectively. We allow  $\nu$  to vary between 0.41 to 0.45 (Hornbach et al. 2004) and  $\alpha$  to range  
300 from 0.67 to 0.77 corresponding to 25% to 0% bulk hydrate in sediments, respectively  
301 (Helgerud et al., 1999).  $P_{water}$  is assumed to be hydrostatic as sediment porosities of ODP  
302 Hole 986C and Hole 986D show a continuous decrease from around 60% near the seafloor to  
303 less than 30% at around 960 mbsf depth (Jansen et al., 1996).  $P_{water}$  is calculated as:

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$$305 \quad P_{water} = \rho_{water}(d_{seabed} + z)g \quad (5)$$

306

307 where  $\rho_{water}$  is assumed to be 1030 kg/m<sup>3</sup> and  $d_{seabed}$  is the water depth below mean sea  
308 level as obtained from the seismic data set.  $P_{gas}$  is calculated at the top of a gas column in a  
309 modelled trap or at the BSR as to:

$$310 \quad P_{gas} = z_{gas}(\rho_{water} - \rho_{gas})g \quad (6)$$

311

312 where  $z_{gas}$  is the modelled gas column height in m and  $\rho_{gas}$  is the associated modelled gas  
313 density in kg/m<sup>3</sup> obtained from the petroleum systems model.

314 For the Mohr-Coulomb-Griffith failure criterion evaluation we assume a dip angle for normal  
315 faults of  $\theta = 60^\circ$  and calculate the coefficients of sliding friction  $\mu = \tan(90-2\theta) = 0.577$  and  $\mu$   
316  $= 0.85$  (Byerlee's law). For the reactivation of faults, we assume no cohesive strength ( $C = 0$ ).  
317 During rock faulting we infer a cohesion of  $C = 2$  MPa for normal faulting and a tensile  
318 strength of  $T_0 = -1$  MPa assuming compacted weak rock types (silty shales and silty  
319 sandstones). Effective stresses ( $\sigma'$ ) are calculated by subtracting the cumulative fluid  
320 pressures from the estimated stress values ( $\sigma$ ) at a given depth.

## 321 4 Results

### 322 4.1 Petroleum systems modelling

323 In this section we will present the results of the hydrocarbon generation and migration  
324 modelling for three different simulation setups. The first setup assumes no GHSZ whilst the  
325 second and third setup consider a GHSZ with gas compositions of scenario A and B as  
326 outlined in Tab. S1 (i.e., assuming 100% methane hydrates or a mix of measured higher  
327 order hydrocarbons in hydrates). At first, we summarize the results common to all three  
328 model setups. Thereafter, we focus on the key differences among the cases and use the  
329 non-GHSZ case as reference frame. In the non-GHSZ case the modelled hydrocarbon  
330 migration and trapping patterns are not affected by methane hydrates and thus reflect the  
331 likely patterns solely governed by the structural and lithological inventory of the model.

### 332 4.2 Similarities in the results of all GHSZ scenario runs

333 The modelling results show that hydrocarbon generation and migration started in the late  
334 Miocene at  $\sim 6$  Ma, from a kitchen area in the AOI<sub>east</sub> being located between the Vestnesa  
335 Ridge in the east and the MTF in the west (Fig. 5). In the AOI<sub>east</sub>,  $213 \times 10^9$  Sm<sup>3</sup> of gas and  
336  $169 \times 10^6$  Sm<sup>3</sup> of oil are modelled to be expelled cumulatively over time. Gas is the dominant

337 hydrocarbon phase modelled to be released from the source rocks and subsequently  
338 migrated, trapped, spilled, and leaked across the basin model. Generated and released oil (>  
339 C<sub>6</sub>) volumes are of minor magnitude and play a secondary role in the modelled hydrocarbon  
340 migration and leakage pattern.

341 With the progressive burial over time, hydrocarbons generated in the western kitchen  
342 migrate exclusively within the Miocene source rock stratigraphic layer in northern, western,  
343 and southern directions toward the physical model boundaries (Fig. 5). In contrast,  
344 hydrocarbons generated in the eastern kitchen migrate within different stratigraphic layers  
345 and in an eastward direction towards the crestal area of the Vestnesa Ridge (Fig. 5, lowest  
346 unit). During time interval 4.37 Ma to 3.2 Ma, eastward directed gas migration intensifies  
347 both within the younger source rock unit (15.18 – 12.05 Ma) and a middle Pliocene  
348 sedimentary sequence (4.37 – 3.2 Ma). After 3.2 Ma, lateral hydrocarbon migration shifts  
349 from the middle Pliocene into a sandier late Pliocene unit (3.2 – 2.7 Ma, Fig. 5). Within this  
350 late Pliocene unit, gas migration is directed towards the Vestnesa Ridge's crest where it  
351 prevails until today (Figs. 5 and 6). This lateral gas migration charges gas accumulations  
352 underneath the crest seeping gas by capillary leakage since ~3 Ma. Over time, gas traps at  
353 shallower stratigraphic levels form within two early Pleistocene age units: eP<sub>A</sub> (~2.7 – 2.1  
354 Ma) and eP<sub>B</sub>, (~2.1 – 1.5 Ma) (Fig. 6).

355 Between 3.2 Ma and 0 Ma, all three model scenarios show an increase in the average  
356 capillary gas leakage rates across all layers and traps from 10<sup>8</sup> to 10<sup>11</sup> Sm<sup>3</sup> Ma<sup>-1</sup>. This equates  
357 to gas flux estimates in the magnitudes 10<sup>1</sup> to 10<sup>4</sup> mg m<sup>-2</sup> day<sup>-1</sup> or 10<sup>-2</sup> to 10<sup>1</sup> kg m<sup>-2</sup> a<sup>-1</sup> (Fig.  
358 7). This gas flux increase is likely due to the progressive eastward maturation of the Miocene  
359 source rocks and thus more proximal supply of hydrocarbons below the Vestnesa Ridge.

360 Through time, many of the crestal gas traps hosted in the Pliocene-Pleistocene units are  
361 modelled to lose gas by (A) capillary leakage, even to the seafloor, and (B) along spill routes  
362 that broadly parallel the Vestnesa Ridge trend. All scenarios show an increase in the  
363 capillary leakage rates to the seafloor from  $10^8$  to  $10^{11}$   $\text{Sm}^3 \text{Ma}^{-1}$  (Fig. 8) from 3.2 to 1.5 Ma,  
364 respectively. The modelled long-distance spill routes may charge other traps and/or exit the  
365 physical boundaries of the AOI<sub>east</sub> (Fig. 6). In all three scenarios, trap charge by spill is  
366 modelled to persist between 3.2 Ma and 0 Ma at average rates of  $2 - 5 \times 10^9 \text{Sm}^3 \text{Ma}^{-1}$  with  
367 maximum values reaching up to  $155 \times 10^9 \text{Sm}^3 \text{Ma}^{-1}$ . Taking the Lunde and Lomvi pockmark  
368 sites as the reference point, a north to north-west directed spill is modelled from traps  
369 located to the west of them. In contrast, a south-east directed spill is modelled from traps  
370 situated to the east of the pockmarks (Fig. 6).

371 Fault set 1 (Fig. 4) is cross cutting hydrocarbon migration routes in the late Pliocene carrier  
372 unit (3.2 – 2.7 Ma) which is charging gas to the Vestnesa Ridge since ~3 Ma (Figs. 5 and 6).  
373 However, the results show that the faults of set 1 are not chosen by the simulator as  
374 preferred leakage pathways into shallower stratigraphic levels. Instead, the hydrocarbons  
375 are modelled to migrate across these fault surfaces and continue within the same carrier  
376 unit (3.2 – 2.7 Ma). The reason for this is 2-fold. First, fault set 1 exhibit negligible fault  
377 throw (at detection limit in the seismic resolution) and hence no offset within the late  
378 Pliocene unit is implemented in the model. Consequently, similar pore entry pressures are  
379 present to either side of the faults in this unit. Second, this unit (3.2 – 2.7 Ma) is  
380 characterized by silty/sandy lithologies and hosts lower pore entry pressures than the  
381 shallower shale dominated units. Therefore, the simulator favors the 3.2 – 2.7 Ma unit for  
382 lateral hydrocarbon migration over fault leakage.

383 All three scenarios yield gas accumulations located directly underneath the Lomvi and Lunde  
384 pockmark seeps (Bünz et al., 2012) within the eP<sub>A</sub> (Trap A) and eP<sub>B</sub> (Trap B) units at present  
385 day (Fig. 6, and Figs. 10 – 13). These gas traps are modelled irrespective of the presence of a  
386 GHSZ (cf. Figs 11 – 13). They are stratigraphic traps that form underneath the ridge's crest  
387 whereby the depth of the accumulations is governed by the lithological distribution (sand-  
388 shale-carbonate) projected from ODP Hole 986D (Jansen et al., 1996). Both traps display a  
389 distinct NW-SE elongation (up to a few kilometers in length) and strongly parallel the  
390 Vestnesa Ridge crestal trend. Among the two units, unit eP<sub>A</sub> hosts the largest amounts of  
391 gas and a major trap, Trap A, is filled to its maximum capacity since ~1.5 Ma. Gas migration  
392 and trapping within both units are modelled to occur at different sublayer levels. Hence  
393 both traps consist of several stacked sublayer accumulations which may have experienced  
394 independent charge-loss histories (Figs. 9 – 13).

#### 395 4.2.1 Key differences in the results of the GSHZ scenario runs

396 Though all three model scenarios yield gas accumulations in the eP<sub>A</sub> and eP<sub>B</sub> units  
397 underneath the seep sites at Vestnesa Ridge since ~2.1 Ma to 1.5 Ma, distinct differences  
398 exist between the scenarios on how much gas is transferred into, lost from, and retained  
399 within both early Pleistocene age units. In the *no GHSZ scenario*, gases leak persistently to  
400 the seafloor until present day at rates between 10<sup>8</sup> to 10<sup>10</sup> Sm<sup>3</sup> Ma<sup>-1</sup> (Figs. 7 and 8). The  
401 maximum seafloor leakage rate occurs at 2.1 Ma (Fig. 7) and coincides with the first overall  
402 peak in capillary leakage (Fig. 8). At this time step, gases accumulate in eP<sub>A</sub> unit in shallow  
403 buried traps, which can sustain only minor gas columns and hence lose gas to the seafloor.  
404 With progressive burial, traps in the eP<sub>A</sub> and eP<sub>B</sub> units accumulate larger column heights and  
405 capillary leakage rates to the seafloor are reduced but leakage remains (Fig. 8).

406 Both *GHSZ model scenarios A and B* show, that upon gas arrival underneath the paleo  
407 Vestnesa Ridge at  $\sim 3.2$  Ma, the inferred GHSZ acts as lateral migration barrier deflecting the  
408 gas migration routes into deeper sublayers. This causes gas trapping and migration at levels,  
409 where, compared to the *no GHSZ scenario*, slightly higher gas columns can be attained  
410 whilst trap capacities are equally small. Therefore, an increased gas spill occurs in the *GHSZ*  
411 *model scenarios A and B* even though the bulk capillary leakage patterns remain similar for  
412 all three scenarios at rates of  $10^8 \text{ Sm}^3 \text{ Ma}^{-1}$  (Figs. 7 and 8). With progressive burial, gas  
413 retention in deeper seated sublayer traps intensifies and more gas accumulates in  $eP_A$  and  
414  $eP_B$  units over time.

415 Hence, the main effect of the presence of a GHSZ in the hydrocarbon migration and trapping  
416 model is to re-route and channel gas migration into the silty-sandy lithologies of the  $eP_A$  and  
417  $eP_B$  units, while in the no GHSZ model scenario gas is constantly leaking upward to the  
418 (paleo-)seafloor. In the GHSZ model scenarios upward seafloor leakage is prevented by  
419 either spill toward the  $AOI_{\text{east}}$  model boundary or by storage in multi-sublayer traps, such as  
420 Trap A (Fig. 9) from -1.5 Ma and -0.148 Ma for scenarios B and A, respectively.

421 Of all three model scenarios, scenario B with the best present-day GHSZ fit (Tab. S1), yields  
422 the strongest confinement of gas within the  $eP_A$  and  $eP_B$  units underneath the Vestnesa  
423 Ridge. In this scenario, Traps A and B accumulate the largest gas volumes, yield the highest  
424 column heights, and are composed of the largest number of sublayer accumulations at  
425 present day (Figs. 9 – 13). Trap A is most prominent underneath the Lunde and Lomvi  
426 pockmarks and is consistently charged at average rates of  $2 - 5 \times 10^9 \text{ Sm}^3 \text{ Ma}^{-1}$  ( $0.6 - 1.4 \text{ kg}$   
427  $\text{m}^{-2} \text{ year}^{-1}$ ) since  $\sim 1.5$  Ma. Trap A is modelled to host  $10.7 \times 10^9 \text{ Sm}^3$  of gas, distributed over  
428 14 sublayer accumulations at present day. Trap B (Figs. 9 – 13) is formed at  $\sim 0.15$  Ma ago

429 and charged by capillary leakage from the  $eP_A$  unit at average rates of  $0.5 - 1 \times 10^9 \text{ Sm}^3 \text{ Ma}^{-1}$   
430 ( $0.14 - 0.3 \text{ kg/m}^2/\text{year}$ ). At present day, Trap B is modelled to host  $0.4 \times 10^9 \text{ Sm}^3$  within two  
431 sublayer accumulation with maximum gas column of 10 m. The shallowest sublayer  
432 accumulation of Trap B is in direct contact with the GHSZ along the crest of the Vestnesa  
433 Ridge (Fig. 13).

#### 434 4.3 Local stress analysis for GHSZ scenario B

435 Analysis of the local stress situation on top of traps in units  $eP_A$  and  $eP_B$  and below the BSR  
436 are shown in Fig. 14. Stress state estimates are based on model values obtained from the  
437 GHSZ scenario B and are given in Tab. 3. At time step -1.5 Ma, a maximum gas column of 25  
438 m is modelled for stacked sublayer accumulations of Trap A whilst Trap B has not been  
439 formed yet. At present day, maximum gas column heights have increased to 46 m and 10 m  
440 in Trap A and Trap B, respectively. These column height values, together with the modelled  
441 corresponding gas density values, are used in the Mohr diagram to evaluate the influence of  
442 gas column pressure on the local stress states (Fig. 14.) Only minor gas amounts are  
443 modelled to accumulate very localized underneath the GHSZ in upper sublayers of the  $eP_B$   
444 unit (Fig. 13). Therefore, the local stress state at the base of the GHSZ is plotted without a  
445 free gas column (Fig. 14) but stress states assuming up to 10 m of gas, alike Trap B, are  
446 calculated and quoted in Tab. 3.

447 All trap top points are modelled to be in stable state of local effective stress as none of the  
448 stress circles are reaching the Coulomb failure criterion (green line in Fig. 14) and hence do  
449 not permit normal faulting nor tensile fracturing at either -1.5 Ma or present day (Fig. 14).  
450 This means that the gas columns of the traps modelled within  $eP_A$  and  $eP_B$  unit underneath  
451 the Vestnesa Ridge's crest do not exert enough pressure to damage intact rocks at any

452 model time. For Poisson's ratios of  $\nu \geq 0.45$ , the gas column pressure is even insufficient to  
453 trigger reactivation by normal faulting, illustrated by the fact that none of the effective  
454 stress scenarios reaches the failure criterion on pre-existing faults (dashed circles in Fig. 14).  
455 However, when assuming Poisson ratio's of  $\nu \leq 0.41$  (i.e., rocks that are more compressible)  
456 the local state of effective stress at both trap top points as well as the base of the GHSZ is  
457 decreased (i.e., the Mohr circles reach the fault reactivation criteria, Fig. 14) for all model  
458 time steps. This means that the modelled gas height can affect the weakened rocks by  
459 triggering normal faulting on pre-existing faults but only for fault planes dipping at  $\geq \sim 60^\circ$   
460 (Fig. 14). The relation between this possible stress state and the plane orientation of fault  
461 set 2 cross cutting Trap A (Fig. 10) and Trap B will be discussed in the following section.

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473 Table 3: Model results used for local stress stability.

Structure	Trap A	Trap B	at BSR
Hosting units (age base – top in Ma)	eP <sub>A</sub> (2.7 – 2.1)	eP <sub>B</sub> (2.1 – 1.5)	eP <sub>B</sub> (2.1 – 1.5)
<i>At 0 Ma model time step</i>			
Water depth (MSL in m)	-1216	-1216	-1222
Burial depth (MSL in m)	-1614	-1421	-1424
$P_{gas}$ (MPa)	0.421	0.092	– (0.092)*
$\sigma'_v$ (MPa)	2.976	1.457	1.569
$\sigma'_H$ (MPa)	0.894 – 1.736	-0.031 – 0.626	-0.182 – 0.525 (-0.245 – 0.445)*
Gas column (m)	46	10	– (10)*
Gas density (kg m <sup>-3</sup> )	96.2648	94.2632	– (94.2632)*
<i>At -1.5 Ma model time step</i>			
Hosting units (age base – top in Ma)	eP <sub>A</sub> (2.7 – 2.1)	–	eP <sub>A</sub> – eP <sub>B</sub> (2.7 – 1.5)
Water depth (MSL in m)	-1236	–	-1241
Burial depth (MSL in m)	-1445	–	-1445
$P_{gas}$ (MPa)	0.229	–	–
$\sigma'_v$ (MPa)	1.401	–	1.630
$\sigma'_h$ (MPa)	-0,067 – 0.526	–	-0.159 – 0.564
Gas column (m)	25	–	–
Gas density (kg m <sup>-3</sup> )	96.5552	–	–

474  $\sigma'$  are effective stresses and account for the influence of pore fluids (water and gas phases).  
 475 Horizontal effective stress ranges ( $\sigma'_h$ ) correspond to Poisson's ratio ranges of  $\nu = 0.41 – 0.45$ . Gas  
 476 densities are derived from the trap information calculated by the Migri simulator. The water depth  
 477 values between above the traps and BSR vary as spatial locations with maximum column heights  
 478 differ from the locations where the GHSZ extents downward into the units. \*) Stress state  
 479 parameters at the BSR are calculated for an assumed 10 m gas column underlying the base of the  
 480 model GHSZ. This setting is chosen as a comparison option for the shallowest sublayer gas traps in  
 481 the eP<sub>B</sub> unit, which are partially overlain by the model GHSZ though the GHSZ is not the primary trap  
 482 seal (Fig. 13).

483

## 484 5 Discussion

### 485 5.1 Thermogenic gas accumulations – the primary source of active methane seepage

486 Hydrocarbon migration modelling results show that accumulation of gas is possible in  
487 dedicated traps underneath the Vestnesa Ridge. Gas started to accumulate in late Pliocene  
488 units, ca. 3 Ma ago, while persistent gas traps (Trap A and B) sustaining seepage at the Lund  
489 and Lomvi pockmarks at present day since the early Pleistocene (~2 Ma) (Fig. 6).  
490 Furthermore, the gas trapped since ~1.5 Ma is confined to structures that are situated  
491 distinctly deeper than the modelled base of the concurrent GHSZ and its BSR expression  
492 (Scenario B; Tab. S1) at present-day (Fig. 13). Below the Vestnesa Ridge, the petroleum  
493 systems model suggests that the presence of a GHSZ (Scenario B) deflects and channels gas  
494 into early Pleistocene units rather than having a sealing effect for substantial amounts of gas  
495 at its base. This corroborates observations by Hushoft et al (2009) along the Svalex P11  
496 seismic line in the AOI (Fig. 1) and Goswami et al. (2015), which show an overall, albeit  
497 outside the gas chimneys, low pore space saturation (less than 30% of the porosity) with  
498 hydrates within the GHSZ and low free gas contents (less than 5%) directly underneath the  
499 BSR. Across the ridge, results of Hushoft et al. (2009) point towards a discontinuous BSR in  
500 fine-grained clay lithologies with higher hydrate saturations in the pore space (up to 11%)  
501 near the vents at the crestal area and lesser values (up to ~6% at the base of GHSZ) away  
502 from chimneys and faults.

503 Based on the opening history of the Fram Strait (Engen et al., 2008), it is reasonable to  
504 assume that the AOI<sub>east</sub> has been at water depths of  $\geq 1000\text{m}$  for the last ~3 Ma.

505 Consequently, the upper strata have been permanently exposed to the temperature and  
506 pressure conditions suited for gas hydrate formation (Sloan and Koh, 2007). This implies

507 that any thermogenic gas migrating at low rates through these strata would have likely  
508 transformed into hydrates, given an appropriate pore water salinity and gas-pore water  
509 dissolution rate are present (Darnell and Flemings, 2015; Spangenberg et al., ~~2005,~~2015).  
510 Modelled capillary leakage fluxes to the paleo-seafloor layer between 3.2 Ma and 1.5 Ma  
511 range from  $\sim 0.05 \text{ kg m}^{-2} \text{ year}^{-1}$  up to  $10 \text{ kg m}^{-2} \text{ year}^{-1}$  ( $\sim 10^1 - 10^4 \text{ mg m}^{-2} \text{ day}^{-1}$ , Fig. 8) for the  
512 GHSZ (Scenario B). Such flux rates are comparable to natural micro- to macro-seepage flux  
513 rates (Etiope and Klusman, 2010) where free, buoyancy driven gas movement occurs from  
514 underground hydrocarbon reservoirs to the earth's surface (Brown, 2000, ~~and~~ Etiope and  
515 Klusman, 2002 and Etiope et al., 2006). At 3.2 Ma, low leakage fluxes are modelled to pass  
516 through shaley lithologies and are captured within the concurrent GHSZ (Liu and Flemings,  
517 2007). For the model times steps from 2.7 Ma to 1.5 Ma, capillary leakage fluxes correspond  
518 to gas flux magnitudes of  $\sim 1 \text{ kg m}^{-2} \text{ year}^{-1}$  that could enable gas chimney formation in the  
519 silty-sandy lithologies (Liu and Flemings, 2006 and Liu and Flemings, 2007) of eP<sub>A</sub> (2.7-2.1  
520 Ma) and eP<sub>B</sub> (2.1-1.5 Ma) units. However, paleo-pockmarks are rather restricted to the time  
521 interval between 1.5 Ma to present (Plaza-Faverola et al., 2015) implying that gas charges  
522 between 2.7 Ma and 2.1 Ma are captured in gas hydrates while the highest modelled gas  
523 fluxes through silty-sandy layers at  $\sim 1.5 \text{ Ma}$  could indicate an initial chimney formation  
524 event and the onset of seafloor seepage (Fig. 8). Seepage periodicity through gas chimneys  
525 or pipes along continental margins has been documented from various settings and using  
526 diverse approaches (e.g., Hovland, and Sommerville 1985; Baraza and Ercilla 1996; Chand  
527 et al., 2012; Davy et al., 2010; Plaza-Faverola et al., 2011; Riboulot et al., 2014). Along the  
528 Vestnesa Ridge vertical fluid migration features (i.e., gas chimneys) are characterized by  
529 highly fractured strata with high amplitude anomalies concentrated at specific stratigraphic  
530 intervals (Fig. 15). High amplitude anomalies are usually present inside seafloor pockmarks

531 (e.g., Panieri et al., 2017) as well as in buried cavities interpreted as paleo-pockmarks (Fig.  
532 15). When the sedimentary layers are not entirely masked by shallow gas accumulations,  
533 high-resolution 3D P-Cable seismic data reveal the presence of these paleo pockmarks and  
534 mounds in the form of erosional surfaces with thicker sedimentary infill (Fig. 15). These  
535 seismic observations have been conceptually explained by the formation of authigenic  
536 carbonate concretions during active seepage events and further burial during less intense  
537 seepage events (e.g., Plaza-Faverola et al., 2011; Plaza-Faverola et al., 2015). Recently,  
538 Himmler et al. (2019) showed a correlation between interpreted paleo features and  
539 authigenic carbonate in deep sediment cores from the Lunde pockmark on the Vestnesa  
540 Ridge. The results of this study validate the conceptual model of buried seepage features  
541 and supporting the inference of an increased number of paleo-seepage events, along the  
542 Vestnesa Ridge, from ca. 1.5 Ma (Plaza-Faverola et al., 2015).

543 Gas migration and trapping models in GHSZ Scenario B show no capillary leakage from traps  
544 to the paleo-seafloor layers for model time steps younger than 1.5 Ma (Fig. 8). This implies  
545 that the multiple methane seepage events recorded during late Pleistocene times at the  
546 Lunde and Lomvi seep sites (Himmler et al., 2019 and Schneider, 2018) are not sourced by  
547 gas leakage due to an overcome of the capillary seal capacities (i.e. seal displacement  
548 pressure sensu Watts, 1987) of traps A and B. Both traps constitute the primary  
549 thermogenic gas sources for the documented seafloor seepage between the late  
550 Pleistocene and today, yet a driving mechanism is required to transfer thermogenic gases  
551 from both traps towards the seafloor and to bypass the GHSZ. We argue that the steeply  
552 dipping faults of set 2 crosscutting Trap A (Fig. 10) presumably constitute the most potential  
553 pathways for gas migration across the GHSZ and thus fueling the late Pleistocene methane  
554 seepage events. We exclude an excess of pore fluid pressure as a cause of mechanical cap

555 rock failure (i.e. hydraulic fracturing, Watts, 1987) above Traps A and B for two reasons. (I)  
556 The depth – porosity trend given for reference site ODP 986 (Jansen et al., 1996) shows a  
557 continuous porosity decrease with depth and provides no indication of deviating  
558 compaction patterns prone to fluid overpressure build up. (II) Local states of stress below  
559 the GHSZ and for Traps A and B do not exert sufficient pore-fluid pressure to trigger  
560 mechanical rock failure of intact rocks (Fig. 14). Modelled gas column heights of up to 46 m  
561 in Trap A (Tab. 3) may permit normal fault reactivation for pre-existing faults with dip angles  
562 larger than  $\sim 60^\circ$  ( $\nu = 0.41$ , Fig. 14) yet require faults to be in place. We, therefore, advocate  
563 tectonically controlled faulting as the most likely governing mechanism to initiate gas  
564 leakage from Trap A since  $\sim 1.5$  Ma and propose a plausible conceptual model of the leakage  
565 system below.

## 566 5.2 Extensional faulting and episodic seepage since $\sim 1.5$ Ma

567 A temporal correlation between seepage events and key glacial periods in the region have  
568 motivated cross-disciplinary studies of the effect of glacial dynamics on the Vestnesa Ridge's  
569 seepage system (Plaza-Faverola et al., 2015, Schneider et al., 2018, Plaza-Faverola and  
570 Keiding 2019, Himmler et al. 2019). It has been suggested that the dynamics at the SBIS  
571 glacial forebulge and/or GIA may have been geological drivers of episodic methane release  
572 along the Vestnesa Ridge (Plaza-Faverola and Keiding 2019; Himmler et al., 2019). Although  
573 uncertainties exist on the extent and volume of circum-Arctic ice sheet beyond the last  
574 glacial period, there is sufficient evidence for the last glacial period that Vestnesa Ridge was  
575 always located to the northwest of the Svalbard-Barents Sea ice sheet (SBIS) margin (Patton  
576 et al. 2015, 2016); most likely in a position where a glacial forebulge would be expected to  
577 form in response to SBIS build-up (Fjeldskaar, 1994). Thus, potential impacts of glacial

578 cyclicities underneath the ice shield on the thermal regime, hydrocarbon generation, phase  
579 composition and migration routes as inferred e.g., on the southwestern Barents Sea shelf  
580 (Ostanin et al., 2017) are considered less likely for this study area on the upper slope. The  
581 extensional faults of set 2 are terminating and crosscutting Trap A located directly  
582 underneath the ridge's crest (Fig. 4 and Fig. 10) and coincide with a position where a glacial  
583 forebulge would be expected to form in response to SBIS build-up (Fjeldskaar, 1994).  
584 Forebulge location, amplitude and bulging response time to glacial loading are sensitive to  
585 the mantle's viscosity as a function of the mantle's temperature (Whitehouse, 2018).  
586 Fjeldskaar (1994) argued for low upper mantle viscosities in the order  $10^{19}$  Pa s<sup>-1</sup> and rapid  
587 mantle relaxation times of a few thousand years during GIA of the Fennoscandian mainland  
588 after the last glaciation. High heat flow values in the study area linked to the nearby mid-  
589 ocean ridge system (Crane et al., 2001; Plaza-Faverola et al., 2017; Hushoft et al., 2009)  
590 suggests that a low asthenosphere viscosity and associated rapid mantle GIA relaxation  
591 times (in the order of decades) are reasonable to assume for the SBIS after the LGM as well.  
592 The post-LGM Fennoscandian forebulge was pinned in its location (~100km from the ice  
593 margin) and responded to the waning ice sheet by continuous reduction in its amplitude  
594 and curvature (Fjeldskaar, 1994). For the SBIS and the Vestnesa Ridge site, we hypothesize  
595 that the glacial forebulge evolution was rather locally fixed and repeatedly changed  
596 amplitude and curvature as a crustal flexure in response to glacial cyclicity (Patton et al.  
597 20105, 20106). Moreover, these responses may have happened within distinctly less than  
598 few thousands of years and even in non-linear fashion (Whitehouse, 2018).

599 Crustal flexure in a forebulge causes minimum horizontal stresses (tensile stresses) in the  
600 upper crust and are orientated perpendicular to the bulge's axis and the ice sheet limits  
601 (Keiding et al., 2015, Lund, 2015). Various modelling approaches on the LGM history of the

602 SBIS show persistent glacial boundaries along the NW-SE running shelf edge, similar to the  
603 trend of the Vestnesa Ridge (Schmidt et al 2014, Patton et al., 2016~~5~~, 2016~~7~~, Fjeldskaar and  
604 Amantov, 2018), and the extensional faults of set 2 (Plaza-Faverola and Keiding, 2019). Lund  
605 et al. (2009) show that during the Weichselian glacial cycle, these minimum horizontal  
606 stresses within the glacial forebulge may attain values up to -10 MPa in the upper 2.5 km of  
607 the crust. Assuming this stress regime applies for the SBIS and the area offshore western  
608 Svalbard, we obtain a reduction in  $\sigma_h$  and a rapid shift of the Mohr circles in the Mohr  
609 diagram to cross the tensile failure criterion (Figs. 14 and 16), however, not the Coulomb  
610 failure criterion for normal faulting (Fig. 14). This would result in the formation of near  
611 vertically dipping faults with no detectable offset and a ~NW-SE strike orientation paralleling  
612 both to the western Svalbard margin and the south-eastern Vestnesa Ridge trend,  
613 analogous to the observed faults of set 2 (Fig. 4).

614 Extensive glacial loading on Svalbard is well documented since the first shelf-edge  
615 glaciations at ~1.5 Ma (Laberg et al., 2010, Mattingsdal et al., 2014) and could have  
616 provoked extensional faulting at Vestnesa ridge's crestal area. We therefore argue that the  
617 faults of set 2 constitute such extensional faults that could have developed in response to  
618 forebulge formation over the Vestnesa Ridge during periods of intense cooling and  
619 extensive SBIS build-up at ~1.5 Ma. Hence, we suggest that the formation of the faults of set  
620 2 as consequence of the first shelf-edge glaciation promoted the primary damage of Trap A  
621 and led to the formation of leakage pathways to the paleo-seafloor.

622 Multiple, intense methane seepage events occurred during four time intervals of intense  
623 cooling and SBIS build-up: (1) during the Penultimate Glacial Maximum (PGM) between  
624 0.16-0.13 Ma (Himmler et al, 2019, Jakobsson et al., 2016, Alexanderson et al., 2018), (2)

625 between 0.05 and 0.04 Ma including Heinrich events H5 and H4 (Himmler et al., 2019,  
626 Kindler et al. 2014, Andrews and Voelker, 2018), (3) Heinrich Event H2 (0.025 Ma) (Ambrose  
627 et al., 2015, Schneider et al., 2018) and (4), Heinrich Event H1 (0.017 Ma) (Szybor and  
628 Rasmussen, 2017, Schneider et al., 2018). Yet, some correlative matches with late  
629 Pleistocene to Holocene interglacial events rather than intense cooling periods are  
630 documented at  $\sim 0.014$  Ma, and  $\sim 0.010 - 0.005$  Ma as well (Himmler et al., 2019 and  
631 Schneider et al., 2018). Still, the distinct episodic nature of the recorded seepage events  
632 during extreme glacial periods clearly suggests that one of the major forces controlling  
633 leakage is associated with extensional faulting at the forebulge. All three migration  
634 scenarios yielded persistent gas charges into both the early Pleistocene age eP<sub>A</sub> unit and  
635 model Trap A between the model timesteps 1.5 Ma and present day. This might indicate  
636 that episodic nature of the seepage events was not governed by a charge fluctuation to the  
637 primary gas reservoirs but rather modulated by episodic faulting and/or changes in the  
638 transmissibility of the fault pathways as a consequence of extensive glaciation on Svalbard.  
639 Intermittent interglacial methane seepage episodes may involve reactivation or dilation  
640 (i.e., pore fluid pressure increase and subsequent negative effective stresses) at the set2  
641 faults (Figs. 14 and 16) induced by tectonic forcing derived from the mid-ocean ridges (e.g.,  
642 Plaza-Faverola and Keiding et al., 2019).

## 643 6 Implications and conclusions

644 The dynamics of Arctic methane seepage in the eastern Fram Strait appear to be governed  
645 by fault re-activation linked to glacio-isostatic adjustments (GIA) due to advancing and  
646 retreating circum-Arctic ice sheets during the Pleistocene (Cremiere et al., 2016, Himmler et  
647 al., 2019). The results of the modelled petroleum system are consistent with the notion of

648 repeated forebulge uplift and subsidence during glacial cycles and associated extensional  
649 fracturing of fault segments at the Vestnesa Ridge since the first shelf edge glaciation of the  
650 Svalbard-Barents Sea ice sheet, ~1.5 million years ago. This fault damage affected the  
651 modelled shallow hydrocarbon reservoirs. This led to the formation of hydrocarbon  
652 migration pathways to the seafloor and repeated tapping of the shallow gas accumulations  
653 ultimately caused the widespread seepage.

654 Similarly, shallow thermogenic gas reservoirs have been confirmed beneath glaciated  
655 continental margins across the circum-Arctic. An area of 33 million km<sup>2</sup> of confirmed  
656 hydrocarbon reserves was directly influenced by grounded ice sheets during the last  
657 glaciation (LG) (Gautier et al., 2009, Andreassen et al., 2017). Thus, glacially driven forebulge  
658 formation too was extensive and ice wastage during numerous Pleistocene glacial-  
659 interglacial cycles have potentially caused large-scale methane release upon deglaciation  
660 due to fracture formation and fault re-activation. Moreover, seafloor seepage linked to  
661 hydrocarbon loss from deeper seated thermogenic hydrocarbons traps during episodes of  
662 deglaciation, erosion, and uplift (Ostanin, et al., 2017) buttress ice shield dynamics as a  
663 major driver forcing thermogenic hydrocarbon release from geological storages. Given the  
664 relatively widespread occurrence of large-scale fracture systems on formerly glaciated  
665 continental margins in the circum-Arctic, it remains to be seen whether methane release  
666 from these systems have attained the atmosphere. Recent analysis of ice cores from  
667 Antarctica points towards a minor contribution of geological methane to the global carbon  
668 inventory during the last deglaciation (Dyonisius et al., 2020). We note, however, that gas  
669 emissions on Vestnesa Ridge is not equivalent to its original old carbon source signal, but  
670 rather biodegraded due to microbial methane formation (Pape et al., 2019, Sauer et al.  
671 2021<sup>9</sup>). A clear source identification of seeping gas in the atmosphere is therefore difficult.

672 More investigations are needed on this topic to explore all the controlling factors of abrupt  
673 natural methane emissions, including re-activation of faults, gas hydrate dissociation and  
674 biodegradation that allow methane emissions at the seafloor and eventually to the  
675 atmosphere.

676 Either way, the 3D petroleum systems model at the Vestnesa Ridge gas hydrate system  
677 reveals that a key process sustaining Arctic methane seepage is the repeated tapping of  
678 hydrocarbon reservoirs due to (a) fracture formation promoted likely by glacial isostatic  
679 adjustments (GIA) and (b) fracture re-activation of existing fault patterns. Our model results  
680 show that migrating and capillary leaking hydrocarbons underneath the Vestnesa Ridge  
681 started to reach the palaeo-GHSZ between 3.2 and 1.5 Ma ago and were focused along-  
682 strike the crestal trend of the paleo-ridge. This implies gas hydrate formation on Vestnesa  
683 Ridge as early as ~3 Ma ago. However, we do not find evidence that substantial amounts of  
684 gas were accumulated beneath the base of the GHSZ. Rather, the gas is trapped in silty-  
685 sandy lithologies of early Pleistocene age and occurs directly below the currently active  
686 Lunde and Lomvi seepage sites where they are crosscut by steeply dipping faults. These  
687 traps are inferred to constitute the most likely thermogenic gas reservoirs sourcing the  
688 episodic seafloor seepage events documented since the first shelf edge glaciation of the  
689 Svalbard-Barents Sea ice sheet, ~1.5 million years ago.

690 Local stress analysis suggests that crosscutting faults underneath the seep site do not  
691 originate from rock fracturing by excess pore fluid pressures. Maximum gas column heights  
692 up to 46 m in the modelled traps are too low to trigger neither normal nor extensional  
693 rock faulting. Hence, we suggest tectonic stresses as the main geological controlling factor,  
694 modulating both faulting activity and gas transmission to the seafloor seepage system. We

695 propose that episodic methane seepage is particularly well coupled to extensional faulting  
696 corrupting the modelled thermogenic gas traps underneath the Vestnesa Ridge crest due to  
697 forebulge development during repeated Svalbard Barents Sea shelf edge glaciations since  
698 ~1.5 Ma.

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711 interpretation of depth horizons, seismic chrono-stratigraphic constraints, and fault surfaces from  
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