A Large West Antarctic Ice Sheet Explains Early Neogene Sea-Level Amplitude

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Early to Middle Miocene sea-level oscillations of approximately 40-60 m estimated from far-field records[1,2,3] are interpreted to reflect the loss of virtually all East Antarctic ice during peak warmth[2]. This contrasts with ice-sheet model experiments suggesting most terrestrial ice in East Antarctica was retained even during the warmest intervals of the Middle Miocene[4,5]. Data and model outputs can be reconciled if a large West Antarctic Ice Sheet (WAIS) existed and expanded across most of the outer continental shelf during the Early Miocene, accounting for maximum ice-sheet volumes. Here, we provide the earliest geological evidence proving large WAIS expansions occurred during the Early Miocene (~17.72-17.40 Ma). Geochemical and petrographic data show glacimarine sediments recovered at International Ocean Discovery Program (IODP) Site U1521 in the central Ross Sea derive from West Antarctica, requiring the presence of a WAIS covering most of the Ross Sea continental shelf. Seismic, lithological and palynological data reveal the intermittent proximity of grounded ice to Site U1521. The erosion rate calculated from this sediment package greatly exceeds the long-term mean, implying rapid erosion of West Antarctica. This interval therefore captures a key step in the genesis of a marine-based WAIS and a tipping point in Antarctic ice-sheet evolution.

Introduction

Reconstructing past Antarctic ice sheet change informs predictions of the continent's contribution to future sea-level rise[6,7]. Since the 1970s, drilling efforts proximal to Antarctica have revealed the general Cenozoic evolution of Antarctic glaciation[8,9,10,11], but fundamental steps in the development of the ice sheets remain poorly constrained. One key uncertainty is the timing of West Antarctic Ice Sheet (WAIS) initiation and expansion across the outer continental shelf. Deep-sea benthic foraminifer oxygen isotope records and Antarctic abyssal plain sedimentary sequences suggested WAIS formation occurred in the Late Miocene or early Pliocene[12,13]. However, drilling from the Antarctic margin[11,14,15] and ice-sheet modelling studies[4,5,16] have raised the possibility that WAIS expansions into areas below sea-level could have occurred during the Early Miocene or earlier, facilitated by a subaerial West Antarctic topography[17,18].

Without widespread WAIS expansions across the continental shelf in the Early Miocene, maximum ice volumes are low enough that global sea-level fluctuations of ~40-60 m estimated from far-field stratigraphic records[1] and oxygen isotope-derived ice volume estimates[2,3] require the near complete loss of the East Antarctic Ice Sheet (EAIS) during the warmest Middle Miocene periods[2]. Such an outcome is incompatible with current ice-sheet model outputs, which suggest retention of most terrestrial East Antarctic ice even during the warmest feasible Middle Miocene environmental conditions[4]. This is mainly due to hysteresis effects driven by height-mass balance feedbacks; once the ice sheet is present, parts of it can be retained in a climate warmer than that which would permit ice-sheet inception on an ice-free landscape[4,19].

Marine sediments, deposited on the continental shelf of the Ross Sea, can reveal whether the WAIS expanded across the continental shelf during the Early Miocene. However, ice proximal geological records have been hampered by poor recovery, unconformities, and/or influence from East Antarctica[9,10,11]. Seismic data suggest that significant volumes of lower Miocene glacimarine sediment exist around the West Antarctic margin[20,21,22,23]. However, seismic data require constraints from drilling to determine the age of the sediments, and to differentiate between detritus from continental-scale ice-sheet expansion and local ice caps on (paleo)topographic highs[22,23]. Consequently, WAIS grounding across the Ross Sea shelf is only clear in seismic data after the Middle Miocene Climate Transition (~14 Ma)[24,25]; it remains uncertain whether there were earlier WAIS expansions across the Ross Sea shelf.

IODP Site U1521 and Provenance Approach

IODP Expedition 374 Site U1521 (75°41.0’ S, 179°40.3’ W; 562 m water depth) was drilled to 650.1 metres below sea floor (mbsf) in the Pennell Basin on the outer continental shelf of the Ross Sea (Fig. 1). The site was drilled in a region that ice-sheet models indicate is one of the last locations where ice grounds during glacial maxima, making it an ideal location to assess the timing of past WAIS expansions onto the outer continental shelf[4,16,26]. The sediments from the base of the borehole up to 209.17 mbsf are split into four chronostratigraphic sequences (1-4; Fig. 2) which constitute an expanded lower Miocene section (~18 to ~16.3 Ma; see Supplementary Material for details) with 73% recovery. These sediments provide a unique window for detailed analysis of ice-sheet behaviour immediately before the onset of the Miocene Climate Optimum (MCO, ~17 Ma; Fig. 2; Extended Data Fig. 1; Extended Data Table 1).

Site U1521 sediments below 209.17 mbsf are predominantly muddy to sandy diamictites, often interbedded with thin laminae and beds of mudstone (see Supplementary Material for details)[26]. Palynological counts on 23 samples revealed sparse palynomorphs in Sequences 1 and 4A, common reworked dinoflagellate cysts in Sequence 2, and evidence for high biological productivity in Sequence 3B (Extended Data Fig. 2; Supplementary Material). These lithological and paleontological data from Sequences 1, 2, 3A and 4A indicate an ice-proximal glacimarine (and potentially subglacial) setting, while data from Sequence 3B suggest an ice-distal setting. Notably, the ~190 m thick succession of Sequence 2, containing a high proportion of reworked dinoflagellate cysts, was deposited rapidly (0.592 mm a^{-1}) within a ~317 kyr interval spanning ~17.72-17.40 Ma (Extended Data Fig. 1).

Through comparison with terrestrial rock outcrops, the sediments recovered at Site U1521 were traced back to their source regions. A differing geological history of the rocks beneath the EAIS and WAIS (Fig. 1) gives the sediment eroded by each ice sheet a distinct geochemical, petrological and mineralogical composition, allowing expansions of the EAIS and WAIS to be distinguished. To avoid bias towards, or omission of, any lithologies, we applied multiple sediment provenance proxies[27]. Specifically, we analysed the detrital fine fraction of 37 samples for neodymium (Nd) and strontium (Sr) isotope compositions (<63 µm) and 23 samples for clay mineralogy (<2 µm). Eight samples were also processed for U-Pb dating of detrital zircons (<300 µm) and five for ^{40}Ar/^{39}Ar dating of detrital hornblende grains (150-300 µm). Additionally, the petrological composition of 15,740 clasts >2 mm was identified down-core (Extended Data Fig. 3).

Evidence for Early Miocene WAIS Growth

At Site U1521, detrital ε_{Nd} values are consistently more radiogenic (higher) in Sequence 2 compared to the sediments above and below (Fig. 2e), implying a contribution from a more radiogenic end member. This end member can be traced to beneath the WAIS; the ε_{Nd} values, ranging between -7.2 and -5.9, are in good agreement with measurements of Upper Quaternary diamicts from the eastern Ross Sea shelf, adjacent to West Antarctica[28]. Here, the radiogenic end member is hypothesised to be the Cenozoic alkali volcanic rocks of Marie Byrd Land, West Antarctica (Extended Data Fig. 4)[28]. Subaerial outcrops of the Marie Byrd Land volcanic province are limited, but magnetic and gravity anomalies associated with subglacial cone-shaped structures indicate the presence of numerous subglacial volcanoes (Fig. 1)[29]. We hypothesize that the Marie Byrd Land volcanic province is the more radiogenic end member in Sequence 2. Conversely, the less radiogenic (lower) ε_{Nd} values seen in Sequences 1, 3A and 4A reflect a mixture of lithologies present in the (East Antarctic) Transantarctic Mountains and fall within the range of Upper Quaternary Ross Sea tills of Transantarctic Mountain provenance (Extended Data Figs. 4, 5)[28,30]. These less radiogenic sediments also show higher and more variable magnetic susceptibility (Fig. 2)[26]. The patterns seen in the ε_{Nd} data are broadly mirrored by detrital Sr isotope compositions (Extended Data Fig. 2).

Single-grain geochronology/thermochronology and clast petrography provide insights into specific source terranes. In the Transantarctic Mountains, Precambrian rocks were affected by the pervasive Ross Orogeny (615-470 Ma), which was accompanied by intrusive felsic magmatism (Supplementary Material)[31]. Zircon age populations from Sequences 1, 3A and 4A show a strong peak towards the earlier part of the Ross Orogeny (595 to 535 Ma) and a 6 to 21% population of Archaean and Paleoproterozoic (>1600 Ma) zircon grains (Figs. 1, 3). These features, together with a lack of grains younger than 250 Ma, resemble data from moraines in the Transantarctic Mountains[32,33,34]. Clasts in sequences 1, 3A and 4A also correlate with rocks in the Transantarctic Mountains, with lithologies including common felsic granitoids and meta-sediments alongside rarer limestones, marbles and sandstones (Extended Data Fig. 3)[31]. Although a relatively minor component, dolerite clasts are found throughout Sequences 1, 3A and 4A (Fig. 2g) and can be traced to the Jurassic Ferrar Group, which predominantly crops out in the Transantarctic Mountains (Fig. 1). Furthermore, rare *Protohaploxypinus* pollen, a distinctive component of the Permian Beacon Supergroup in the Transantarctic Mountains, are observed in Sequence 3A[35]. Overall, the sediments comprising Site U1521 Sequences 1, 3A and 4A are predominantly sourced from erosion of the Transantarctic Mountains in East Antarctica.

In contrast, Sequence 2 is characterized by the highest ε_{Nd} values and contains zircons with Cretaceous (~100 Ma) U-Pb ages (n = 16; Fig. 3a). Such ages are indicative of a West Antarctic provenance as they are presently only found beneath the modern Siple Coast ice streams, including Kamb Ice Stream and those closer to Marie Byrd Land[33,36]. The age spectra of samples from Sequence 2 share other features with data from the Siple Coast ice streams, including a broad Triassic (~240-190 Ma) age peak, few pre-Mesoproterozoic zircons (<5 % of grains) and a young (~515-505 Ma) Ross Orogeny peak (Fig. 3)[33]. Detrital hornblende $^{40}Ar/^{39}Ar$ ages from Sequence 2 further corroborate a West Antarctic provenance. Unlike zircon grains, which can survive multiple sedimentary cycles, hornblende grains are less resistant to weathering. The absence of Grenvillian (~1100-900 Ma) ages in the Sequence 2 hornblende sample (Extended Data Fig. 6) therefore suggests a West Antarctic provenance, as Grenville-age rocks are absent there[37]. The scarcity of Ferrar Group dolerite clasts, common in the Transantarctic Mountains, is also consistent with a West Antarctic provenance (Figs. 1, 2), as is a high proportion of smectite in the clay fraction at the bottom of Sequence 2 (≤58%; Extended Data Fig. 3), with smectite percentages similar to Quaternary sediments in the eastern Ross Sea[38]. Additionally, Sequence 2 contains evidence for recycling of older marine detritus, most likely from the lower Cenozoic rift-fill strata that exist in the eastern Ross Sea region of the West Antarctic Rift System[21]. This is inferred from the dominance of reworked Eocene-Oligocene species in the diatom and spore-pollen assemblages[26], alongside the common (13-21%) reworked Eocene-Oligocene marine dinocysts, which are rare (<1.5%) in younger sediments (Extended Data Fig. 2).

Smectite abundance declines significantly up-section within Sequence 2 and is accompanied by an increase in the proportion of basalt clasts (Extended Data Fig. 7). This anticorrelation is unexpected given that smectite is a weathering product of basalt and volcanic rocks. We infer that lower in Sequence 2, basaltic bedrock was predominantly weathered to smectite and was thus largely confined to the finer grain size fractions. Over time, this more weathered regolith layer was removed, leading to erosion of progressively more pristine continental detritus containing more basalt clasts. This scenario is supported by more radiogenic ε_{Nd} values measured in the <63 µm fraction lower in Sequence 2 (Fig. 2, Extended Data Fig. 7), as Marie Byrd Land basalts are more radiogenic than other lithologies (Extended Data Fig. 5). Sequence 2 (17.72-17.40 Ma) could therefore record an advance of the WAIS over parts of West Antarctica which had not been covered by grounded ice for an extended period.

Further evidence for WAIS expansion can be found in seismic data, which can trace the sediment package deposited at Site U1521 between 17.72 and 17.40 Ma (Sequence 2) across the Ross Sea continental shelf[23]. The sediment package, which is thicker towards the eastern Ross Sea (i.e., West Antarctica), contains glacial features including widespread progradational wedges and high relief morainal banks[20,21,23]. Coupled with the lithological and palynological evidence for ice proximity at Site U1521, this shows marine-terminating ice was present. Transport of large volumes of West Antarctic detritus as far west as the Pennell Basin in the central Ross Sea is evident in our provenance data, which, alongside common reworked marine microfossils, proves this marine-terminating ice derived from an Early Miocene WAIS which intermittently extended across most of the outer continental shelf.

Our data therefore reveal WAIS expansions across the Ross Sea continental shelf date back to at least 17.72 Ma, which is significantly earlier than previously suggested[12,13,23,24,39]. Advance of the WAIS into marine-based areas (i.e., regions grounded mainly below sea level) at 17.72-17.40 Ma is supported by a corresponding period of high sensitivity of the marine δ^{18}O record to obliquity forcing (Fig. 2i). High obliquity sensitivity is considered a proxy for enhanced ice-sheet sensitivity to ocean dynamics and thus the presence of marine-based ice[15].

Birth of a Marine-Based WAIS

The mean erosion rate for the WAIS catchments draining to the Ross Sea between 17.72 and 17.40 Ma can be estimated using the volume of the corresponding seismic package east of Site U1521[23]. Assuming that, at the time of deposition, the area of the Ross Sea drainage sector of the WAIS was approximately the same as today (±20%), the inferred sediment volume requires a mean catchment erosion of approximately 87 m in ~317 kyr (Extended Data Table 2). The mean erosion rate of ~0.275 mm a^{-1} during this interval greatly exceeds the long-term mean rate of 0.012 mm a^{-1} calculated for this part of the WAIS between 23 and 14 Ma[18]; even when the full uncertainty is taken into account (Extended Data Table 2), it is still more than an order of magnitude higher. This highlights the 17.72 to 17.40 Ma interval as one of unusually rapid erosion, with erosion rates comparable to modern subpolar to temperate glacial catchments[40]. Transporting this large volume of subglacially eroded debris quickly to the WAIS margin required abundant meltwater at the ice sheet bed[41], as well as fast-flowing ice streams that extended into marine settings where broad deposition took place. Ocean temperatures must therefore have been sufficiently cool to permit the advance of marine-based ice, yet atmospheric conditions must have remained warm enough to provide sufficient precipitation to drive dynamic ice flow and enhanced basal erosion[4].

Since most of West Antarctica, apart from Marie Byrd Land, was thermally subsiding throughout the Miocene[18], the high erosion rate at 17.72 to 17.40 Ma is unlikely to have been driven by tectonic uplift. The eroded sediments therefore reflect ice expansion and enhanced glacial incision of the terrestrial West Antarctic hinterland, plus infilling of the Ross Sea basins. This erosive event occurred at a time when topographic reconstructions indicate a transition from a terrestrial West Antarctic topography (23 Ma) to a largely submarine West Antarctic topography (14 Ma)[18]. The timing and large volume of sediment deposited in Sequence 2 at Site U1521 suggests that the 17.72 to 17.40 Ma interval records a critical step in the transition of the WAIS from a largely terrestrial ice sheet to one that was primarily marine-based. This significant alteration to West Antarctic topography occurred just prior to major changes affecting the Antarctic cryosphere and global climate during the MCO[2,11]. Subglacial erosion may therefore have driven changes in ice-sheet evolution and behaviour as, after ~17.40 Ma, a greater submarine area in central West Antarctica would have made the mass-balance control of the WAIS more sensitive to external drivers such as sea level and oceanic forcing[5,16]. We propose that ice retreat at the onset of the MCO may be partially attributable to the crossing of this topographic tipping point and that Sequence 2 records the birth of a marine-based WAIS. This event dates to well before 14 Ma, the time slice at which topographic reconstructions first show a largely sub-marine West Antarctica[18].

Sea-Level Reconciliation

Grounded ice flowing from West Antarctica was close to Site U1521 towards the end of the Early Miocene. We therefore validate recent modelling studies suggesting that an ice sheet nucleating on a partially terrestrial West Antarctica could expand extensively into the marine realm under Early Miocene climatic and paleotopographic conditions[4,5,16]. Our data are consistent with an ice extent similar to, or exceeding, the largest modelled Early to Middle Miocene Antarctic ice sheets (Fig. 1), containing ice volumes of approximately 80 m sea-level equivalent (SLE) depending on the topographic reconstruction used[4,5,16]. This expanded WAIS contained approximately 14-15 m SLE of ice, but also acted to buttress the EAIS resulting in significantly larger-than-present ice volumes[4,16]. These maximum ice volume constraints indicate that far-field sea-level amplitudes of ~40-60 m did not require the loss of nearly all terrestrial East Antarctic ice during subsequent warm periods during the MCO[1,2,3], consistent with modelled EAIS hysteresis effects[4]. By providing the earliest conclusive evidence for a large marine-based WAIS, our data also dispel long-held inferences that a WAIS, able to significantly impact global eustacy and climate, was not present until the Middle or Late Miocene[12,13,39].

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Figure Legends

Figure 1. Site U1521 location and surrounding geology. The outcropping regional geology around the Ross Sea[42] (© SCAR GeoMAP and GNS Science 2019) is overlain on the BedMachine Antarctica V1 modern bed topography[43,44]. The MEaSUREs grounding line, ice sheet margin and basins are used[45,46] and the map was produced using ArcGIS software. IODP Site U1521 is located on the outer continental shelf of the central Ross Sea. Locations referenced in the text are labelled, including the

ANDRILL 2A (AND-2A) and Cape Roberts Project 1 (CRP-1) drill sites. The white dashed line indicates the boundary between East and West Antarctic lithosphere[47]. Orange triangles show Cenozoic subglacial volcanic edifices detected based on morphological characteristics, gravity anomalies and magnetic anomalies[29]. The inset shows an ice-sheet model simulation using a 'cold' climate ('cold' orbit and a climate with 280 ppm atmospheric CO_2 concentrations) and an estimated Middle Miocene topography[4]. Provenance indicators from Site U1521 Sequence 2 sediments are broadly consistent with an ice sheet similar to or exceeding the extent of this model output.

Figure 2. Selected provenance proxies from IODP Site U1521 compared to Early Miocene climate records. The light blue shaded section (Sequence 2) highlights the interval with sediments of predominantly West Antarctic provenance. The depth of Ross Sea Unconformity (RSU) 4a and 5 and seismic surface D-b are indicated in red[23]. a) Site U1521 inclination data after 20 mT demagnetisation (red points)[26] and polarity interpretation (white = reverse polarity, black = normal polarity, grey = no interpretation). b) Site U1521 lithostratigraphy[26]. c) Chronostratigraphic sequences. The circled letters between b) and c) mark the depths of the zircon U-Pb samples (Figure 3). d) Magnetic susceptibility measured on the whole core[26]. e) Neodymium isotope signature of the fine fraction. Error bars are 2 S.D. external reproducibility; for provenance interpretations, see Extended Data Figure 4 and references in Supplementary Material. f) Abundance of Eocene-Oligocene dinocysts as a percentage (black) and concentration (i.e., counts per gram sediment; grey). g) Dolerite clast abundance. Errors shown in f) and g) are 95% confidence intervals[48]. Magnetostratigraphic tie points between the polarity interpretations from shipboard data (a)[26] and geomagnetic polarity timescale (h)[49] are marked by purple dashed lines. i) Obliquity sensitivity, indicating the strength of obliquity in the $\delta^{18}O$ record relative to the theoretical strength of obliquity forcing. This has been interpreted as representing the presence of marine-based Antarctic ice[15]. j) Sea-level record based on an oxygen isotope splice[2]. Red and blue shaded intervals indicate pronounced sea-level highstands (>40 m) and lowstands (<-20 m), respectively. MCO = Miocene

Climatic Optimum. k) CO_2 reconstruction with a LOESS smoothing (shaded region indicates 1 sigma error)[50]. l) Simplified lithological log from the AND-2A record, with diamictites differentiated based on a grounding-zone proximal vs distal glacimarine depositional setting[11,15].

Figure 3. Site U1521 detrital zircon U-Pb age distributions. a) Data displayed as kernel density estimates (KDEs). When present, large Ross Orogeny (~600-500 Ma), Triassic (~240-190 Ma) and Cretaceous (~100 Ma) age peaks are labelled. The age ranges of the Ross Orogeny, Grenville Orogeny and a ~2.7 Ga event recorded in Ross Sea sedimentary strata are illustrated using grey-shaded bars. The sub-bottom depth midpoints of the samples are shown in Figure 2 and listed in the methods section. b) Same data as in a), displayed as a multi-dimensional scaling (MDS) map calculated using the Kolmogorov–Smirnov statistic[51]. Stress (a measurement of the goodness of fit between the disparities and the fitted distances[51]) = 0.072. A MDS plot visualises the degree of similarity between samples, with the proximity of sample points reflecting their similarity. The axis scales are dimensionless and have no physical meaning. The colour of Site U1521 samples (A to I) corresponds their ε_{Nd} value. Previously published zircon U-Pb data from Kamb, Whillans and Bindschadler ice streams in West Antarctica, as well as Transantarctic Mountain moraines from inland and coastal regions, are shown in grey[32,33,34]. The KDEs and region of the MDS plot interpreted as having a West Antarctic provenance are shaded in light blue, consistent with the blue shading in panel a) and Figure 2. Note that although Whillans Ice Stream drains the WAIS, it is excluded from the blue shaded area due to its proximity to the Transantarctic Mountains (Figure 1), resulting in a subglacial sediment provenance signature indistinguishable from East Antarctic detritus[33].

Methods

Neodymium and Strontium Isotopes

Samples were disaggregated and wet sieved to isolate the <63 µm fraction, which was then dried at 60°C. This size fraction represents the bulk composition, as samarium and neodymium are incorporated in equal proportions into most rock-forming minerals, meaning grain-size sorting is not likely to impact results[52,53]. However, the Rb-Sr system is subject to elemental fractionation during weathering and grain-size sorting, which can influence ^{87}Sr/^{86}Sr ratios (see 'Provenance Changes within Sequence 2' section in Supplementary Material). To remove authigenic Fe-Mn oxyhydroxide phases, samples were leached in a mixture of 0.05 M hydroxylamine hydrochloride, 15% acetic acid, and 0.03 M EDTA at a pH of 4[54]. A carbonate removal step was not included due to the very low carbonate content[26]. Leached sediment was dried, homogenised, and 50 mg aliquots were digested on a hotplate in concentrated HF (2 mL), $HClO_4$ (0.8 mL) and HNO_3 (1 mL) for three to five days, with a subsequent 6 M HCl step. The Nd was isolated from the sample matrix using a cation exchange resin (AG50W-X8, 200-400 µm mesh) and HCl in increasing molarity, followed by a low molarity HCl Ln-Spec resin procedure (50–100 µm mesh). The sample matrix from the cation exchange step was dried down, taken up in HNO_3, then loaded onto Eichrom Sr Spec resin to wash down the matrix and elute the Sr[55].

Neodymium isotopes were measured in the MAGIC laboratories at Imperial College London on a Nu high resolution multi-collector inductively coupled plasma mass spectrometer (HR MC-ICP-MS). To account for instrumental mass bias, isotope ratios were corrected using an exponential law and a ^{146}Nd/^{144}Nd ratio of 0.7219. Although negligible, interference of ^{144}Sm on ^{144}Nd was corrected for. Bracketing standards were used to correct measured ^{143}Nd/^{144}Nd ratios to the commonly used JNdi-1 value of 0.512115[56]. USGS BCR-2 rock standard was processed alongside all samples and yielded ^{143}Nd/^{144}Nd ratios consistently within error of the published ratio of 0.512638 ± 0.000015[57]. Full procedural blanks for Nd ranged from 7 to 30 pg (n = 6). ^{143}Nd/^{144}Nd ratios are expressed using epsilon notation (ε_{Nd}), which denotes the deviation of a measured ratio from the modern Chondritic Uniform Reservoir (0.512638)[58] in parts per 10,000.

Strontium isotopes were measured in the MAGIC laboratories at Imperial College London on a TIMS (Thermal Ionisation Mass Spectrometer). 10% of the sample was loaded in 1 µL of 6M HCl onto degassed tungsten filaments with 1 µL of $TaCl_5$ activator. The measured $^{87}Sr/^{86}Sr$ ratios were corrected for instrumental mass bias using an exponential law and an $^{88}Sr/^{86}Sr$ ratio of 8.375. Interference of ^{87}Rb was corrected for using an $^{87}Rb/^{85}Rb$ ratio of 0.386. Analyses of the NIST 987 standard reference material were completed every four unknowns, yielding a mean of 0.710290 ± 0.000041 (2SD, n = 36). Samples were corrected to the published value of 0.710252 ± 0.000013[57]. The relatively poor reproducibility for our NIST 987 runs was due to technical issues, but is still more than sufficient for interpreting sample results, which change in the 3rd to 4th digit. Accuracy of results was confirmed using rock standard USGS BCR-2, processed with every batch of samples, which yielded $^{87}Sr/^{86}Sr$ ratios of 0.705010 ± 0.00029 (2SD, n = 18). This is well within error of the published ratio of 0.705013 ± 0.00010[57].

Detrital Zircon U-Pb Dating

The sub-bottom depth midpoints of the nine samples are: A: 220.23, B: 270.03, C: 335.72, D: 373.58, E: 410.82, F: 487.40, G: 546.55, H: 588.00 and I: 642.21 mbsf. To ensure there were enough grains for statistical analysis, the above samples were taken over 40 cm intervals. Samples were disaggregated, dried and sieved at 300 µm. Zircons from the <300 µm fraction were concentrated using standard gravity settling and magnetic separation techniques. Samples were then mounted in resin, polished and analysed using an Agilent 7900 laser ablation inductively-coupled plasma mass spectrometer (LA-ICP-MS) with a 25-35 µm pit diameter in the London Geochronology Centre at University College London. Approximately 150 grains resembling zircons were randomly selected for analysis from each sample. Plešovice zircon[59] was used as a primary standard to correct for instrumental mass bias and depth-dependent inter-element fractionation. Approximate U and Th concentrations were calculated by comparison with NIST 612 glass[60].

Data reduction of the time-resolved mass spectrometer data was performed using GLITTER 4.5(61). Ages younger than 1100 Ma were calculated using the ^{206}Pb/^{238}U ratio whilst older grains used the ^{207}Pb/^{206}Pb ratio. Data were filtered to exclude non-zircons based on zirconium concentrations (>10^6 counts per second) and a -5/+15% discordance threshold was applied. This yielded at least 92 grains per sample, giving a 95% confidence that any age populations comprising more than 7% of the sample will be measured[62]. GJ1 zircon[63] was used as a secondary standard to verify accuracy of the data. Repeat analyses using zircons with and without existing ablation pits were made to check sample reproducibility; these agreed within the uncertainties associated with random sampling. Final data were processed and visualised using the R package IsoplotR[64].

Clast Petrography

The gravel fraction (>2 mm) was characterized in continuum along the core between 648.17 and 209.17 mbsf. Clasts exposed in the cut surface of the archive half core were measured, logged and described on the basis of macroscopic features (e.g. shape, colour, texture). Logging aimed to identify the distribution and variation of the gravel-size clasts along the core length. Clast logging followed the methods previously applied to the ANDRILL and Cape Roberts Project drill records from the Ross Sea. On the basis of macroscopic features, clasts were grouped into seven main lithological groups: igneous rocks, quartz fragments, dolerites, volcanic rocks, metamorphic rocks, sedimentary rocks and sedimentary intraclasts[65,66,67,68]. Data processing involved counting the occurrence of each lithological group over 10 cm core intervals and summarizing this for each core (Extended Data Fig. 3). The total number of clasts was also summed for each metre interval (Extended Data Fig. 3). To highlight the along-core variation in dolerite and volcanic clasts - two of the most indicative lithologies for provenance constraint - the number of these clasts was divided by the total number of clasts in each core (Extended Data Fig. 3). A total of 73 pebble to cobble-sized clasts were sampled for petrographic analysis, of which the most representative of each lithological group were analysed using standard petrographic methods with polarized light microscopy.

Palynology

Sample processing was performed at Utrecht University, following standard techniques of the Laboratory of Palaeobotany and Palynology. Samples were oven-dried and weighed (~15 g dry weight sediment each). One *Lycopodium clavatum* tablet with a known amount of marker spores was added for quantification of palynomorph abundances[69].

Samples were treated with 10% HCl (Hydrochloric acid) and cold 38% HF (Hydrofluoric acid), then sieved over a 10 μm mesh with occasional mild ultrasonic treatment. To avoid any potential processing-related preservation bias, no oxidation or acetolysis was carried out. The processed residue was transferred to microscope slides using glycerine jelly as a mounting medium, and 2 slides were analysed per sample at 400× magnification. Slides were examined for detailed marine palynomorphs (dinoflagellate cysts, acritarchs and other aquatic palynomorphs) and, at screening-level, terrestrial palynomorphs (pollen and spore) at Utrecht University. Subsequent detailed analysis of terrestrial palynomorphs on a sub-set of seven samples was undertaken at GNS Science. Of the 23 palynological samples analysed for dinocysts, two contained <60 dinocysts (Sequence 1; 594.48 mbsf and Sequence 2; 567.75 mbsf) and one was almost barren (yielding only 12 *in situ* dinocysts, Sequence 3A; 374.9 mbsf). The almost barren sample is excluded from all plots. The two low abundance samples are included in our plots but require careful interpretation. Samples between 594.48 and 567.75 mbsf and below 594.48 mbsf (cores 65R, 67R, 69R and 71R) were also checked, but yielded few dinocyst specimens. Those present comprised of fragments of mostly reworked dinocysts.

Pollen and spore identification followed taxonomic compilations[70,71], augmented by key Antarctic literature[72,73,74]. For pollen and spores, scanning continued until an entire cover slide was completed, or a 100 count reached. Results are presented as specimens/gram, and percentage of all terrestrial palynomorphs. Dinocysts were identified based on a taxonomical index[75] and informally and formally described species in the literature[76,77,78,79]. Dinocyst percentages were calculated based on the total *in situ* dinocysts counted, excluding reworked specimens. The percentages of other palynomorph groups such as brackish and freshwater algae (*Cymatiosphaera* spp. and *Pediastrum* spp.) and reworked dinocysts were calculated using the total palynomorphs counted (Fig. 2; Extended Data Fig. 2). *In situ* dinocyst and terrestrial palynomorph absolute abundance (specimens/g dry weight) and the absolute abundance of the other palynomorph groups were calculated by counting the amount of *Lycopodium clavatum* spores encountered, following the equation of Benninghoff (1962)[80].

Protoperidinioid (P) dinocysts are mostly represented by the genera *Brigantedinium*, *Lejeunecysta*, and *Selenopemphix*. Gonyaulacoid (G) dinocysts mostly include *Batiacasphaera* spp., *Operculodinium* spp. and *Spiniferites* spp. Protoperidinioid cyst percentages (Heterotrophic % in Extended Data Fig. 2) and percentages of the most common species (*Brigantedinium* spp. *Lejeunecysta* spp., *Selenopemphix* spp. and *Selenopemphix antarctica*) were calculated to identify productivity trends and/or the presence of sea ice (see Supplementary Material). P dinocysts are likely produced by heterotrophic dinoflagellates[81] and, at present, dominate the assemblages in Antarctic sediments in areas with high nutrients and/or (year-round) sea-ice cover. At present, samples in quasi perennial sea-ice covered areas are dominated by *Selenopemphix antarctica* (~75%), with abundant *Brigantedinium* spp. and rare occurrence of other species[82,83,84]. G cysts are generally produced by phototrophic dinoflagellates. *Operculodinium* spp. is the most abundant, has species representatives among the extant cysts and has been selected to represent temperate-warm conditions. At present, it is almost exclusively found in temperate areas of the Southern Ocean north of the Subantarctic Front and never occurs in circum-Antarctic sediments south of the Polar Front[82]. In contrast, it is common to abundant in other Antarctic warm Miocene records[85,86]. Reworked dinocysts include Eocene and Oligocene taxa (mostly *Vozzhennikovia* spp., but also few *Spinidinium* spp. and *Enneadocysta diktyostila*).

Sediment Volume Estimate

The volume of sediment comprising Sequence 2 was estimated based on seismic data for the Ross Sea continental shelf[23]. The isopach maps were developed by interpolating between available seismic reflection profiles[23], giving a total volume of 175,526 ± 17,553 km^3. The 10% uncertainty accounts for uncertainty in seismic velocities, which vary from 1700-2700 ms^{-1} at Site U1521 based on tomography and 1970-2480 ms^{-1} based on down-hole measurements. As the provenance data suggest a West Antarctic sediment source for Site U1521 Sequence 2, we assume that all the sediments east of 180° and south of 73° are derived from West Antarctica. This is the vast majority (123,627 ± 12,363 km^3) of the sediment across the shelf. Our sediment volume estimate is conservative, as the top of Sequence 2 (surface D-b) has been truncated across much of the continental shelf by RSU4[23]. Significant sediment volumes are also likely to be present beyond the edge of the seismic data from the continental rise. Any sediment beneath the modern Ross Ice Shelf is also unaccounted for, although this component is likely to be small.

To translate this sediment volume into an erosion rate, the approach and uncertainty range of Paxman et al. (2019)[18] was used to account for porosity and a small biogenic sediment component (Extended Data Table 2). We note that using generic values in our porosity calculation is crude, with variation in the porosity of these Antarctic sediments likely to be significant[86], but nevertheless sufficient for our order-of-magnitude estimate of erosion. It is reasonable to assume the major ice divides have remained in largely the same positions since the Early Miocene, as indicated by various modelling studies using reconstructed topographies[4,5,16]. The size of the eastern Ross Sea catchment (i.e. Ross Sea sector of the WAIS) was therefore assumed to be similar to the modern, with a 20% uncertainty. Some sediment in these units clearly contains reworked material; there are high concentrations of

Eocene-Oligocene palynomorphs and diatoms. Although this means our erosion rate is not indicative of pure bedrock incision, it still represents a significant change to the topography and bathymetry of West Antarctica. The material removed likely exceeds our conservative estimate of ~87 m across the catchment. The 317,416 year duration is based on the cyclostratigraphic analyses described in the age model section, with a 20,000 year uncertainty.

IODP Site U1521 Age Model

The age model for IODP Site U1521 uses magnetostratigraphy, biostratigraphy, cyclostratigraphy, ^{87}Sr/^{86}Sr dating of macrofossils, and ^{40}Ar/^{39}Ar ages of hornblende grains to correlate rock units to the Geomagnetic Polarity Timescale (GPTS)[49]. Key events and tie points are summarized in Extended Data Table 1 and illustrated in Extended Data Figures 1 and 8. Biostratigraphic constraints include first and last appearance datums of diatoms. The maximum and minimum age range reported for these datums are derived from total and average ranges[88,89] and hybrid range models derived from Constrained Optimization (CONOP) methods[89,90]. Hybrid range model ages are used as primary constraints for our age model. This is because they best account for up section reworking of microfossil datums, which is common in glacial sedimentary environments, whilst recognising that major down section reworking is unlikely (partly because of the rarity of bioturbated intervals). They are marked by base of arrows in Extended Data Figure 1 and mentioned in the text below. Biostratigraphic datums and magnetic polarity reversals provide tie points to construct lines of correlation (LOC) with the GPTS. The age model presented here includes the interval of West Antarctic sediment provenance (Sequence 2) and is described from the base of the borehole at 650 mbsf to 75 mbsf.

Biostratigraphic constraints through the interval from 650 mbsf to near the top of Sequence 3B (at ~286.1 mbsf) are sparse as the sediments are deeper than the Opal-CT transition and diatom preservation is relatively poor. Our correlation of the four distinct magnetozones R5, N4, R4, and N3 to the GPTS is therefore primarily based on regional correlation of prominent seismic reflectors to other dated drill cores from the Ross Sea shelf, backed up by diatom biostratigraphic constraints. The section from 650 mbsf to 567.95 mbsf at Site U1521 (Sequence 1) is characterised by reversed magnetic polarity but offers no constraints which we can confidently use for correlating this reversed interval to the GPTS. However, robust age constraint for sediments at the base of Sequence 2 can be determined through regional correlation of RSU5 to other sites where chronostratigraphic data are available. RSU5 intersects Site U1521 at 567.95 mbsf (the base of Sequence 2) and is correlated across the Glomar Challenger Basin and tied to DSDP Site 273 at 282 mbsf[23]. The LAD of *T. praefraga* is observed at 309 mbsf in DSDP Site 273, which suggests that RSU5 is younger than 17.95 Ma at that site. RSU5 cannot be directly correlated into the western Ross Sea, but a major unconformity (U2) occurs in the AND-2A drill site at 774.94 mbsf and likely corresponds with RSU5 based on chronostratigraphy[11]. Specifically, sediments that directly underlie U2 in AND-2A are characterised by a reversed magnetic polarity and are correlated to Chron C5Er (18.636 to 18.497 Ma) based on constraints that include ^{40}Ar/^{39}Ar dates of 18.82 ± 0.15 Ma on pumice clasts within a tuffaceous siltstone at 831.66 mbsf. The age of sediments that overlie U2 at the AND-2A drill site are constrained by the FAD of *T. praefraga* at 771.5 mbsf (<18.46 to 18.58 Ma), and a ^{40}Ar/^{39}Ar date of 18.04 ± 0.31 Ma on pumice clasts within a tuffaceous siltstone at 709.17 mbsf. These observations require correlation of the reversed magnetic polarity zone that characterise the sediments above U2 to Chron C5Dr.2r (18.007 to 17.676 Ma). All evidence presented above shows that between ~18.6 and ~17.8 Ma, a significant, regionally extensive, erosional event (or series of events) created surface RSU5/U2.

Sediments deposited on top of RSU5 at Site U1521 are characterised by reversed magnetic polarity. Based on the known age of RSU5 at DSDP Site 273 and U2 at AND-2A, we correlate the top of reversed magnetozone R5 in Site U1521 to Chron C5Dr.2r. This interpretation is consistent with the observation that *T. praefraga* is not present in a diatom-bearing sample at 563 mbsf, despite comprehensive searches for this species in this sample as well as diatom-bearing samples higher in Sequence 2. As *T. praefraga* is a small and compact diatom not prone to fragmentation which would likely be preserved in the observed diatom assemblages, we are confident this absence is not a result of poor preservation below the Opal-CT transition. *T. praefraga* is a common species in upper Oligocene and lower Miocene sediments recovered from several sites across the Ross Sea, including Cape Roberts Project-2/2A, DSDP Site 273, and AND-2A[11,91,92]. The total reported CONOP model based age range for the LAD of *T. praefraga* is 17.95 to 16.82 Ma and the hybrid model range is 17.95 to 17.36 Ma[88,89,90]. Consequently, we view the absence of *T. praefraga* as strong evidence that the sediments above 563 mbsf at Site U1521 are younger than 17.95 Ma.

We then correlate the magnetic polarity reversal (MPR) R5/N5 between 526.8 and 524 mbsf to C5Dr.1n/C5Dr.2r (17.676 Ma), the MPR N4/R4 between 517.2 and 515.1 mbsf to C5Dr.1r/C5Dr.1n (17.634 Ma), and the MPR R4/N3 between 400.5 and 397.2 mbsf to C5Dn/C5Dr.1r (17.466 Ma). We extend a line of correlation from this MPR to the top of Sequence 2, where it intersects with seismic surface D-b[23]. The correlation presented here by interpolating through these MPRs indicates sediments in Sequence 2 span the time interval from ~17.7-17.4 Ma. The occurrence of the diatom taxon *Thalassiosira* sp. cf. *T. bukryi* at 450.52 mbsf supports this correlation as the range reported for this taxon at ODP Site 744 is 17.7-17.4 Ma[89,93].

To refine the likely sedimentation rate and timespan of Sequence 2, a cyclostratigraphic analysis was conducted on clast abundance data (Extended Data Fig. 3) spanning 568 to 380 mbsf. These data were analysed using TimeOpt[94], which is a statistical optimization method for astronomical time scale construction and astrochronologic testing, executed by the *astrochron* package in R[95] (function 'timeOpt'). Given a range of plausible sedimentation rates and a series of specified astronomical periodicities (for precession, obliquity, and eccentricity), TimeOpt identifies the age model that results in a time-series that best aligns with the predictions of Milankovitch theory. Specifically, two diagnostic attributes of the astronomical hypothesis are evaluated: the hierarchy of cyclic frequencies expected of Milankovitch Cycles, $r^2_{spectral}$, and the match between eccentricity cycles and the precession-band envelope, $r^2_{envelope}$[94,96]. These two values (r^2_{power} and $r^2_{envelope}$) are multiplied to produce an r^2_{opt} value, which provides insight into the strength of a hypothesized astronomical signal at each evaluated sedimentation rate.

Assuming plausible average sedimentation rates between 40 cm kyr^{-1} and 65 cm kyr^{-1}, TimeOpt yields an optimal sedimentation rate of 59.2 cm kyr^{-1} for Sequence 2, with an r^2_{opt} of 0.396. To assess the statistical significance of the result, a Monte Carlo astrochronologic test is conducted to evaluate the null hypothesis that the observed variability in clast abundance arises entirely by stochastic processes, rather than astronomical forcing. The Monte Carlo simulations are generated using the function “timeOptSim”, which creates a large number of similar time-series of stochastic (“red”) noise, to assess the probability that such datasets can produce an r^2_{opt} value comparable to the one generated by the clast abundance data[94,96]. This analysis yields a p-value of 0.005, indicating that the null hypothesis (i.e. the data is generated from a stochastic “red noise” process; specifically an AR1 process) can be rejected with a high degree of confidence. Given that the astrochronologically-estimated sedimentation rate is derived independently from the paleomagnetic data, their consistency is remarkable and provides strong evidence in support of an estimated duration of ~317 kyrs for Sequence 2 (Extended Data Fig. 1)[26].

While the ‘floating’ TimeOpt-derived astronomical time scale preserves information about elapsed time, it must be separately anchored to a specific numerical age. To do so, we use the ‘slideCor’ function in the *astrochron* package[95]; this is an automated approach to find the optimal anchoring of the floating TimeOpt-derived time scale to the theoretical astronomical solution of Laskar et al. (2004)[97]. Specifically, we have applied a Taner bandpass filter[98], isolating the periods between 60 ka and 27 ka for both the obliquity component of the astronomical solution[97], and for the TimeOpt-derived floating astrochronology. The optimal match between the astronomical solution and floating astrochronology is identified using the squared Pearson correlation coefficient.

Independent biostratigraphic and magnetostratigraphic constraints mean we can restrict our 'slideCor' assessment to a feasible ~800 ka interval; our lower limit (17.950 Ma) is based on the absence of *T. praefraga* and the correlations of RSU5 described above, and our upper limit is based on the C5Cr/C5Dn MPR (17.154 Ma). Since the precise relationship between clast abundance and astronomical forcing is not known with certainty, any time-anchor for the astronomically calibrated section should be treated as having an uncertainty of at least a full obliquity cycle (~41 ka). Application of the slideCor function identifies two plausible regions of the astronomical solution for anchoring the Sequence 2 clast abundance data. The optimal match ($r^2 = 0.8497$) results in an astronomically calibrated section ranging from 17.601 Ma to 17.918 Ma (± 0.02 Ma). This would indicate that the interval is ~140-220 kyrs older than the age range suggested by the paleomagnetic interpretation, giving a very poor match with the measured polarities. However, a slightly less optimal match ($r^2 = 0.7704$) anchors the section to span 17.398 Ma to 17.715 Ma (± 0.02 Ma), which places it within ~40 kyrs of the paleomagnetic interpretation. This agreement of geochronological frameworks derived from paleomagnetism and astrochronology, which are broadly independent, provides strong support for the age model presented here.

Uncertainties in the magnetostratigraphic age model, most notably for Subchrons C5Dr.1n and C5Dr.1r and Chron C5Dn, may account for some of the slight disagreement with the astrochronlogical approach described above. The available astronomically tuned durations of these (sub-)chrons agree within 10%[99,100]. The small discrepancies in duration of (sub-)chrons originate from the astronomical tuning approach (carbon and oxygen isotopes tuned to eccentricity, tilt and precession at Site 1090(99) and carbonate content to eccentricity only at Site U1336(100)), as well as physical and palaeomagnetic recording processes such as bioturbation and the palaeomagnetic lock-in depth[101,102]. Paleomagnetic measurement methods are discussed in detail in the cruise report[26].

We suggest 17.95-17.40 Ma as the absolute uncertainty of the timing of Sequence 2 deposition, based on the absence of *T. praefraga* (17.95 Ma) and occurrence of MPR C5Dn/C5Dr.1r (17.466 Ma) near the top of Sequence 2. However, more precise constraint on the duration of Sequence 2 deposition can be achieved based on the remarkable agreement of sedimentation rates based on the astronomical analysis of clast data and interpolation through magnetostratigraphic tie points, which suggest deposition occurred over ~317 kyrs. Combined with the close correlation between our astrochronological analyses and the timing of MPRs, we suggest a more precise interval for the deposition of Sequence 2, spanning ~17.72-17.40 ± 0.02 Ma. The ~20 kyr error represents uncertainty in the phase relationship between clast abundance and obliquity forcing. This range coincides closely with many independent records indicating ice-sheet growth, including a sea-level lowstand recorded on the New Jersey continental margin (~17.8-17.46 Ma)[1], evidence for ice sheet growth in the AND-2A drill core sediments (~17.8-17.4 Ma)[11] and a peak in obliquity sensitivity (~17.8-17.5 Ma)[15] (Fig. 2).

The age of Sequence 3A and 3B (324.20- 209.17 mbsf), bracketed by seismic surface D-b and regional unconformity RSU4a, is difficult to tightly constrain. Diatom preservation increases significantly in a sample at 286.1 mbsf at the base of Sequence 4A and the FADs of *Nitzschia sp. 17 Schrader*, *Synedropsis cheethamii*, and *Denticulopsis maccollumii* suggest sediments below this stratigraphic level are older than 17 Ma. The LAD of *F. maleinterpretaria* in this sample provides a minimum age constraint and suggests that the sediments below 286.1 Ma must be older than 16.41 Ma. These constraints require that the sediments between 344.6 and 286.3 mbsf, characterised by reversed polarity, correlate with either the Subchron C5Cn.2r or the base of Chron C5Cr. Correlation to the base of Chron C5Cr is our favoured option as this would indicate that the interval of time missing across seismic surface D-b is relatively short, whereas regional unconformity RSU4a at the top of this unit records a hiatus of longer duration. The alternative interpretation is shown with a dashed line in Extended Data Figure 1.

We constrain the slope of the LOC through Sequence 3B based on the sedimentation rate indicated for the diatom-bearing Sequence 4B as the sediments are similar, although affected by diagenesis in Sequence 3B. The sedimentation rate in Sequence 3A is assumed to be comparable to the Sequence 2 diamicts. We also acknowledge that the actual first appearance of the diatom taxa identified in the sample at 286.1 mbsf may have originally been deeper, but their presence has since been obscured by diagenesis. This would require that the LOC sit to the left (younger) of its current position. Therefore, we include an error box (orange box in Extended Data Fig. 1) in our age model to show that the LOC could occur anywhere within this area depending on the amount of time missing across D-b and the sedimentation rate during deposition. We are confident that the MPR between 400.5 and 397.2 mbsf (N3/R3) is C5Dn/C5Dr.1r (17.466 Ma) based on constraints above and below this interval outlined above and place our LOC through the reversal. This LOC requires a time gap of ~ 180 kyrs across regional seismic surface D-b[23] that separates Sequences 2 and 3.

The relatively thin interval of reversed polarity within Chron C5Dn (at ca. 380 mbsf) is not identified in the current version of the GPTS (Extended Data Fig. 8), but a similar short-duration reversed polarity event roughly halfway through Chron C5Dn is recorded in the AND-2A magnetostratigraphic record[11]. Taking the palaeomagnetic uncertainties of ice-proximal sediments into account, we hypothesise that this rarely recorded reversed polarity event could be a genuine feature of the geomagnetic field that has not been detected in marine sediments due to signal smoothing at low sedimentation rates[103].

The age of sediments above RSU4a are very well constrained by diatom data, ^{87}Sr/^{86}Sr ages and magnetostratigraphy. The LAD of *F. maleinterpretaria* indicates that the sediments above 286.1 mbsf must be younger than 16.41 Ma. An ^{87}Sr/^{86}Sr date on shell fragments at 272.65 mbsf indicates the interval with reversed polarity containing the fragments correlates with Subchron C5Cn.1r (16.351 to 16.261 Ma). This correlation means that the hybrid age model underestimates the maximum age of the FAD of *Nitzschia grossepunctata*, which occurs at 286.1 mbsf, and suggests the age indicated by the total range model age for this datum (16.23 Ma) is more likely. Together, these data indicate that the base of Sequence 4A dates to less than ~16.351 Ma. We correlate the MPR (R3/N2) between 209 and 205 mbsf to C5Cn.1n/C5Cn.1r (16.261 Ma). The sequence of well-dated shells through Sequence 4B allows us to correlate the sediments between 209 and 106.3 mbsf that are characterised by normal polarity with Subchron C5Cn.1n (16.261 to 15.994 Ma) and the MPR between 106.3 and 105.5 to C5Br/C5Cn.1n (15.994 Ma). The FADs of *Denticulopsis lauta*, *Actinocyclus ingens*, *Denticulopsis hyalina*, and *Denticulopsis simonsenii* at 84.99 mbsf indicate a major hiatus at this depth spanning from ~15.83 Ma to at least 14.48 Ma. This stratigraphic horizon correlates with RSU4, a major regional unconformity[23].

Sediment Provenace Interpretations

To interpret the provenace data from IODP Site U1521, they must be placed in a regional context. In the Supplentary Material, we therefore present a short gelogical summary of the Ross Sea secotor[31,37,104-170], including a compilation of published zircon U-Pb data[33,105-129]. We also include a more detailed discussion of our hornblende ^{40}Ar/^{39}Ar[136,154,171-175], clast petrography, clay mineralogy[158,183-188], and palynology[35,72,78,84,189-192] datasets. Additional insights into the sediment proveance of Sequences 1, 2 and 3A are also explored[23,33,38,184,193-199]. A compliation of literature neodymium and strontium isotope data (visualised in Extended Data Figures 4 and 5) is provided in Supplementary Table 1.

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Data availability The datasets generated as part of this study are available in the British Geological Survey National Geoscience Data Centre. Datasets include Nd and Sr isotope data (https://doi.org/10.5285/3a646c8a-8422-4079-a928-a159532439eb), zircon U-Pb dates (https://doi.org/10.5285/cfadf931-0804-484c-a9d0-96254239c421), clast counts (https://doi.org/10.5285/b043471f-22e5-40e4-b274-1c875316d725), clay mineralogy data (https://doi.org/10.5285/b3cb3574-49b0-44c8-a934-3da88ca4ef93), hornblende ^{40}Ar/^{39}Ar dates (https://doi.org/10.5285/926cad28-669f-4703-8a5b-5e7e843a4ee1) and palynological counts (https://doi.org/10.5285/adea0809-5fe5-4fb5-9f3e-9d774534d26d).

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Author Contributions

J.W.M., T.v.d.F, R.M.M., L.D.S. and A.E.S. designed the research in collaboration with the entire IODP Expedition 374 science party. J.W.M. conducted the Nd and Sr isotope analyses. L.Z., F.T. and M.P. performed the clast counts. J.W.M., P.V. and A.C. produced the zircon U-Pb data. F.B. and V.B.R. collected the clay mineralogy data. F.S., J.P. and C.B. performed the palynological counts and interpretations. S.R.H. provided the hornblende ^{40}Ar/^{39}Ar data. K.J.L. provided guidance on geochronology interpretations. L.F.P., F.C. and L.D.S. calculated the sediment volume estimate. R.L, R.M.M., T.E.v.P., D.H., D.K.K. and E.M.G. improved the shipboard age model. N.B.S and S.R.M. conducted the astrochronological analyses. D.K.K. provided the XRF data. E.G. and B.K. helped integrate sediment provenance data with numerical modelling. I.B., G.K., and J.E.D. advised on specific technical aspects of the manuscript. J.W.M. created the figures and wrote the text with assistance from all authors and particular guidance from T.v.d.F, C.D.H, E.G., and M.J.S.. All Expedition 374 scientists contributed to the collection of shipboard datasets and the interpretations of the data.

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Competing Interests The authors declare no competing interests.

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Extended Data Figure and Table Legends

Extended Data Figure 1. Age model constraints below 75 mbsf at Site U1521. From left to right are: depth (metres below sea floor), core number, core recovery (black = recovered), inclination prior to and after 10 and 20 mT demagnetisation (black, blue and red points, successively), and corresponding polarity interpretations (black = normal, white = reversed, grey = no interpretation). Note that the polarity interpretations have been simplified compared to those in the cruise report[26], with small uncertainties related to core gaps removed. Note Site U1521 is in the Southern Hemisphere. The geomagnetic polarity timescale[49] is shown across the top of the plot. The orange shaded regions indicate uncertainties in our age model and the dashed line marks an alternative line of correlation for Sequence 3. The blue line indicates the age model for Sequence 2 based on our astrochronological analyses, with the light blue shading indicating the ~20 kyr uncertainty associated with the phase relationship between clast abundances and obliquity. This astrochronological anchoring agrees closely with linear interpolations between magnetostratigraphic tie points (black line).

Extended Data Figure 2. Selected palynological counts compared to strontium and neodymium isotope data. Palynological data are reported as percentages (crosses) and counts/gram (circles). The blue shaded area represents Sequence 2, which is interpreted as consisting of sediments with a West Antarctic provenance. Error bars indicate a 95% confidence interval[48].

Extended Data Figure 3. Down-core clast and clay mineral distribution. The blue shaded area highlights Sequence 2, which is interpreted to consist of sediments with a West Antarctic provenance. a) Core lithology (see Figure 2 for key). b) Chronostratigraphic sequences. c) Clast abundance. d) Percentages of different clast lithologies. e) Ratio between dolerite and total number of clasts (red) and volcanic rocks and total number of clasts (green), with 95% confidence interval shown as pale shading[48]. f) Clay mineral abundances.

Extended Data Figure 4. Map of approximate ε_{Nd} values in rocks and offshore sediments from around the Ross Sea embayment. Epsilon Nd values are overlain on MODIS imagery[210] and the BedMachine Antarctica V1 modern bed topography[43,44], with the MEaSUREs grounding line and ice sheet margin shown[45,46]. The approximate boundary between West and East Antarctic lithosphere is shown using a white dashed line[47]. Modern/late Holocene and terrestrial till samples are represented by circles with the same colour bar[28,30,55]. Although ice flow patterns have changed since their deposition, Last Glacial Maximum tills in offshore sediments are also plotted as squares to improve spatial coverage[28]. Individual samples and references are reported in Supplementary Table 1. The bedrock map was produced by Kriging between sample locations within a group, then masking to the outcrop area. Beacon and Ferrar Group (Fig. 1) rocks are often not differentiated in geological mapping, but are roughly equal volumetrically[136], with the uppermost Beacon Supergroup formations having a Ferrar-like isotopic signature[139]. We hence assume a 60% Ferrar, 40% Beacon mixture is representative.

Extended Data Figure 5. Kernel density estimate plots for literature measurements of rock ε_{Nd} compared to measurements on fine-grained Miocene detritus from Site U1521. For references and a list of all the data, see Supplementary Table 1. The height of the curve indicates the density of measurements and n the total number of samples analysed. Colour scheme is identical to Figure 1, with sediments in grey.

Extended Data Figure 6. Kernel density estimates for hornblende $^{40}Ar/^{39}Ar$ ages compared to zircon U-Pb ages younger than 1500 Ma. The two dating methods are show in red and blue, respectively. Bold letters correspond with those in Figure 3. The positions of major peaks and number of grains analysed are labelled in the corresponding colours. Stratigraphic position is shown in Figure 2.

Extended Data Figure 7. Close up of the Site U1521 interval with a West Antarctic provenance. The stratigraphic log (a) is displayed alongside the percentage of reworked dinocysts (b), basalt clast fraction (c), relative abundance of smectite (d), Nd isotope data (e) and Fe/Ti ratios determined by X-ray fluorescence scanning (f).

Extended Data Figure 8. Correlation of Site U1521 magnetostratigraphic tie points. Shown are correlations between the AND-2A record[11], Site U1521[26] and the GPTS[49].

Extended Data Table 1. Age tie points for Site U1521 below 75 mbsf. FAD: First Appearance Datum, LAD: Last Appearance Datum. Depth errors for the biostratigraphic datums reflect the position of the first downhole sample in which the reported species was not observed. We cannot exclude the possibility that the true first observation occurs between this sample and that reported as the FAD. Opal-CT indicates that the lowermost occurrence is uncertain due to poor preservation below the Opal-CT transition (~286.1 mbsf). Age errors for the biostratigraphic events are given as the maximum and minimum reported ages based on hybrid range models[89,90]. Magnetic Polarity Reversals (MPR) depths are given as midpoints between samples with differing polarities, with the depth error indicating the distance to these samples.

Extended Data Table 2. Values used in the erosion rate calculation.