

Scalar Relativistic Effects with Multiwavelets: Implementation and Benchmark

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ABSTRACT: The importance of relativistic effects in quantum chemistry is widely recognized, not only for heavier elements but throughout the periodic table. At the same time, relativistic effects are strongest in the nuclear region, where the description of electrons through a linear combination of atomic orbitals becomes more challenging. Furthermore, the choice of basis sets for heavier elements is limited compared with lighter elements where precise basis sets are available. Thanks to the framework of multiresolution analysis, multiwavelets provide an appealing alternative to overcoming this challenge: they lead to robust error control and adaptive algorithms that automatically refine the basis set description until the desired precision is reached. This allows one to achieve a proper description of the nuclear region. In this work, we extended the multiwavelet-based code MRChem to the scalar zero-order regular approximation framework. We validated our implementation by comparing the total energies for a small set of elements and molecules. To confirm the validity of our implementation, we compared both against a radial numerical code for atoms and the plane-wave-based code EXCITING.



INTRODUCTION

In his famous Nobel Price Lecture, P.A.M. Dirac stated that with the advent of quantum mechanics, all fundamental problems of chemistry were in principle solved.¹ It is however interesting that he did not seem to realize that relativity would also play a role, despite his own fundamental contribution to combining relativity and quantum mechanics.^{2,3} It is now widely accepted that the correct description of the electronic structures of atoms and molecules requires the inclusion of relativistic effects. This is particularly relevant for the core- and the innermost valence shell-electrons of heavier elements, which move at sizable fractions of the speed of light. A wide number of chemical properties and phenomena, such as the yellow color of gold,^{4,5} the liquid state of mercury,^{5,6} the functioning of lead-acid batteries,^{5,7} the catalytic behavior of cobalt,^{5,8–10} and the difficulties to describe fluorinated compounds with density functional theory (DFT),^{11–13} all are due to relativistic effects. There is a wide spectrum of methods that include relativistic effects in electronic structure calculations. They range from the use of effective core potential (ECP),^{14–17} also known as pseudopotentials in the solid-state physics community,¹⁸ to the solution of the four-component Dirac equation. In between these two extremes, there are a number of methods with different degrees of accuracy and complexity.¹⁹

The ECP method collapses a given number of core electrons and the nucleus to provide a core potential for the valence electrons. This has the dual advantage of reducing the number of explicit electrons while at the same time implicitly including relativistic effects. On the other hand, if core electronic properties such as high pressure chemistry, core–electron spectroscopy²⁰ and nuclear magnetic resonance (NMR) shielding constants⁹ are considered, ECP methods fall short, and all-electron calculations are necessary. The four-component Dirac equation² constitutes the starting point for the treatment of relativistic effects, and the full Breit Hamiltonian^{21–25} is the most complete treatment of relativistic effects for a many-electron system. On the other hand, it is also the most computationally expensive: spin–orbitals are four-component complex functions. Spin is no longer a good quantum number in a relativistic framework, and the simplified picture of two electrons with opposite spin sharing the same orbital is no longer valid.^{19,26,27} Most operators couple the

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different components of a spin-orbital, leading roughly to a factor 100 in computational cost, because coupling four complex functions requires 8×8 matrices instead of a real scalar.^{28–31} Although progress has been made to make four-component calculations faster and easily available,^{28–31} it is still convenient to attempt approximations that promise great reduction in the computational cost at the price of reduced accuracy. The first step in such a hierarchy of approximations is the elimination of the small components through a Foldy–Wouthuysens (FWs) transformation.^{32,33} The choice of transformation leads to different kinds of methods, such as the Pauli Hamiltonian,^{34–36} the regular approximations (RAs),^{37–39} the Douglas–Kroll–Hess (DKH) Hamiltonian,^{40–44} or the exact two-component (X2C) Hamiltonian method.^{19,45–53} Further reduction to a scalar method is also possible for the Pauli Hamiltonian and the RAs.^{37–39} In particular, the RAs have two interesting features: (1) the decoupling part of the FWs transformation can be expanded in a convergent series to recover the exact elimination, and (2) the renormalization part can be exactly incorporated into the wave function. The zero-order regular approximation (ZORA)^{34,37,54–56} is the simplest form of Hamiltonian keeping only the zeroth order in both parts. The reduction to a scalar method is carried out by applying the Dirac identity and discarding the spin-orbit term. The advantage of ZORA is to keep most of the standard algorithms of quantum chemistry in their original form just rescaling the kinetic energy by a function that includes the potential energy (see below Theory and Implementation section).^{34,37,54–56} The numerical treatment of this rescaling function can be challenging because it displays a cusp at each nucleus.^{34,37,54–56} Standard approaches, based on atomic orbital expansions, can struggle to obtain an accurate description in this region. The issue is further aggravated for heavier nuclei, where such a correction is important, and at the same time, the number of basis sets available is more limited and less is known about their true precision.^{57,58}

Some efforts to assess the precision of ZORA for all-electron calculations of heavier elements have been undertaken using Gaussian type orbitals (GTOs), but it is anyway challenging to assess the precision without an external reference.^{57,58} One such option for hydrogen-like ions (He^+ , Ne^{9+} , Ar^{17+} , etc.) is constituted by the scaling properties of the ZORA Hamiltonian in a two-component framework, which yields the (scaled) exact Dirac energies for such a system.⁵⁹ However, for many-electron systems and/or for a scalar relativistic approach, assessing the true precision of a GTO basis is challenging, and several basis sets are developed to include relativistic effects. Examples of all-electron relativistic basis sets include the universal Gaussian basis set (UGBS),⁶⁰ the atomic natural orbitals basis sets,^{61–65} the X2C basis sets,⁶⁶ and the segmented all-electron relativistic contracted basis sets.^{67–71} Importantly, such basis sets must be fitted to the chosen Hamiltonian (ZORA and DKH). This comes on top of the required fitting of the basis set to a given electronic structure method, and if relevant, to a particular property.⁷¹ Although some of the known all-electron relativistic basis sets may provide good results for a particular method and property, their transferability is limited. The large number of available GTO basis sets is an indication that no single basis set is good enough to describe all properties of interest to sufficient precision.⁵⁸

In recent years, multiwavelets (MWs)⁷² have emerged as a powerful alternative to traditional local basis sets.⁷³ Their foundation based on multiresolution analysis⁷⁴ leads to a basis set that is not empirically parametrized. Robust error control^{75–77} means that the user can set a finite but arbitrary target precision, and adaptive algorithms^{78–80} ensure that the representation of molecular orbitals is automatically refined until the required precision is reached. Instead of a plethora of bases to choose from, the user only needs to set the requested precision and the polynomial order of the basis, providing in practice a robust black-box, where straightforward numerical considerations guide the user's choice. MWs have proven reliable and robust in providing Hartree–Fock (HF) and DFT benchmark results for energies⁷³ and properties⁸¹ and have ventured both toward post-HF methods⁸² and relativistic treatments.⁸³

In this work, we present a MW implementation of the ZORA method in the MRChem code. To verify the correctness of the implementation, we consider two alternative methods: a radial atomic solver implemented as a separate package⁸⁴ and linearized augmented plane wave (LAPW) implemented in the electronic-structure code EXCITING.^{85,104} Both of these approaches provide means for systematic improvement of the precision and are capable of yielding total energies approaching the complete basis set limit, as demonstrated in refs 84 and 86.

THEORY AND IMPLEMENTATION

From the Dirac Equation to the ZORA Hamiltonian.

We will briefly expose how the ZORA^{34,37,54–56} is derived starting from the four-component Dirac equation of an electron, which describes a four-component spinor Ψ , subject to a potential V

$$(\beta c^2 + V + \alpha \mathbf{p})\Psi = E\Psi \quad (1)$$

In the above equation \mathbf{p} is the momentum operator, c is the speed of light, atomic units ($m_e = 1$, $\hbar = 1$, $e = -1$) are assumed and α and β are defined as follows

$$\beta = \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix}, \quad \alpha = \begin{pmatrix} 0 & \boldsymbol{\sigma} \\ \boldsymbol{\sigma} & 0 \end{pmatrix}$$

$$\sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \sigma_y = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma_z = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad (2)$$

The first step toward ZORA consists in applying the FW transformation to the Dirac Hamiltonian of eq 1

$$H^{\text{FW}} = U^\dagger H^D U \quad (3)$$

where the transformation matrix $U = W_1 W_2$ is a product of a decoupling matrix W_1 and a renormalization matrix W_2

$$W_1 = \begin{pmatrix} 1 & -R^\dagger \\ R & 1 \end{pmatrix}, \quad W_2 = \begin{pmatrix} 1/\sqrt{1+R^\dagger R} & 0 \\ 0 & 1/\sqrt{1+RR^\dagger} \end{pmatrix} \quad (4)$$

and R is the exact coupling between the large and the small components of a four-spin-orbitals

$$R = \frac{1}{2c^2 - V + E} c\boldsymbol{\sigma} \cdot \mathbf{p} \quad (5)$$

The inverse potential term depends on the eigenvalue E , and this dependence can be expanded in a Taylor series as follows

$$\begin{aligned} R &= \frac{1}{2c^2 - V} \left(1 + \frac{E}{2c^2 - V} \right)^{-1} \boldsymbol{\sigma} \cdot \mathbf{p} \\ &= \frac{1}{2c^2 - V} \sum_{k=0}^{\infty} \frac{(-E)^k}{(2c^2 - V)^k} \boldsymbol{\sigma} \cdot \mathbf{p} \end{aligned} \quad (6)$$

Restricting the expansion to the zeroth order leads to RA. Once the renormalization W_2 is also considered, the following two-component Hamiltonian is obtained

$$H^{\text{RA}} = \frac{1}{\sqrt{1 + R^\dagger R}} \left[V + \boldsymbol{\sigma} \cdot \mathbf{p} \frac{c^2}{2c^2 - V} \boldsymbol{\sigma} \cdot \mathbf{p} \right] \frac{1}{\sqrt{1 + R^\dagger R}} \quad (7)$$

The ZORA Hamiltonian is finally obtained by Taylor-expanding the renormalization operator and retaining only the zero-order term

$$H^{\text{ZORA}} = V + \boldsymbol{\sigma} \cdot \mathbf{p} \frac{c^2}{2c^2 - V} \boldsymbol{\sigma} \cdot \mathbf{p} \quad (8)$$

Using the Dirac identity

$$(\boldsymbol{\sigma} \cdot \mathbf{A})(\boldsymbol{\sigma} \cdot \mathbf{B}) = \mathbf{A} \cdot \mathbf{B} + i\boldsymbol{\sigma} \cdot (\mathbf{A} \times \mathbf{B}) \quad (9)$$

we can separate the scalar-relativistic and spin-orbit contributions to the kinetic energy covered by the first and the second terms, respectively. Here, we keep only the scalar-relativistic part and obtain the following Hamiltonian

$$H_{\text{SR}}^{\text{ZORA}} = V + \mathbf{p} \frac{c^2}{2c^2 - V} \cdot \mathbf{p} = V + \frac{1}{2} \boldsymbol{\kappa} \cdot \mathbf{p} \quad (10)$$

where in the last expression, we have implicitly defined $\boldsymbol{\kappa} = \left(1 - \frac{V}{2c^2} \right)^{-1}$. Given the ZORA Hamiltonian, a Kohn-Sham (KS) DFT implementation is then obtained by replacing the nonrelativistic kinetic energy operator with its ZORA counterpart

$$\left(\frac{1}{2} \boldsymbol{\kappa} \cdot \mathbf{p} + V \right) \varphi_i = \epsilon_i \varphi_i \quad (11)$$

It should be noted that in practical implementations, the potential defining $\boldsymbol{\kappa}$ is usually *not* the full KS potential. Given the form of $\boldsymbol{\kappa}$, the most important contribution is nuclear attraction. Introducing Coulombic (\hat{J}) and exchange and correlation (\hat{V}_{xc}) is possible, but the corresponding operator has to be recomputed numerically at each iteration. Additionally, operations such as function multiplications are difficult to perform in traditional linear combination atomic orbital basis representations. A common choice for GTO calculations is to use a fixed atomic potential, which is called the atomic-ZORA approximation.⁸⁷ The ZORA kinetic operator can then be precomputed and used throughout the calculation. Atomic-ZORA has the advantage of being gauge-invariant.⁸⁷

ZORA Equations in a Multiwavelet Framework. To obtain a MW implementation of the ZORA eigenvalue problem, it is necessary to transform the differential equation into an integral equation, in analogy with the nonrelativistic case.^{75,77} The standard KS equations can be concisely written as follows

$$\hat{F} \varphi_i = \sum_j F_{ij} \varphi_j \quad (12)$$

where \hat{F} is the Fock operator, φ_i refer to an occupied molecular orbital, and F_{ij} are the matrix elements of the Fock operator between two occupied orbitals, assuming a general non-canonical (nondiagonal) form.

Within the framework of KS-DFT, the Fock operator consists of the kinetic energy \hat{T} , the nuclear attraction \hat{V}_{nuc} , the Coulombic repulsion \hat{J} , the HF exchange \hat{K} scaled by some numerical factor $\lambda \in [0, 1]$, and the exchange and correlation potential \hat{V}_{xc}

$$\hat{F} = \hat{T} + \hat{V}_{\text{nuc}} + \hat{J} - \lambda \hat{V} + \hat{V}_{\text{xc}} \quad (13)$$

In the nonrelativistic domain, the coupled KS differential (eq 12) can be rewritten in the integral form,^{88,89} by making use of the bound-state Helmholtz kernel

$$\varphi_i = -2\hat{G}_{\mu_i} (\hat{V} \varphi_i - \sum_{j \neq i} F_{ij} \varphi_j) \quad (14)$$

where $\hat{G}_{\mu_i} = [-\nabla^2 - \mu_i^2]^{-1}$ is the integral convolution operator associated with the bound-state Helmholtz kernel $G(r) = e^{-\mu_i r}/r$, using $\mu_i = \sqrt{-2F_{ii}}$ (i.e., the diagonal elements of the Fock matrix).

In the ZORA Hamiltonian, the kinetic energy operator becomes

$$\begin{aligned} \hat{T} &= \mathbf{p} \cdot \frac{c^2}{2c^2 - V_Z} \mathbf{p} = -\frac{c^2}{2c^2 - V_Z} \nabla^2 - \nabla \frac{c^2}{2c^2 - V_Z} \cdot \nabla \\ &= -\frac{1}{2} \boldsymbol{\kappa} \nabla^2 - \frac{1}{2} \nabla \boldsymbol{\kappa} \cdot \nabla \end{aligned} \quad (15)$$

where $V_Z = \hat{V}_{\text{nuc}} + \hat{J} + \hat{V}_{\text{xc}}$. Including \hat{J} and \hat{V}_{xc} does not pose any issue in the MW framework, other than the computational overhead of having to update the potential at every iteration, because all potentials are anyway treated on an equal footing using a numerical grid. The nonlocal HF exchange, in turn, does not seem to contribute significantly to V_Z as shown in ref 90.

Inserting the ZORA kinetic operator into eq 12, we obtain

$$\left(-\frac{1}{2} \boldsymbol{\kappa} \nabla^2 - \frac{1}{2} \nabla \boldsymbol{\kappa} \cdot \nabla + V \right) \varphi_i = \sum_j F_{ij} \varphi_j \quad (16)$$

In order to make use of the same framework as the nonrelativistic implementation of eq 14, it is necessary to isolate the Laplacian and the diagonal element of the sum on the right-hand side, which together make up the bound-state Helmholtz operator $\hat{G} = (\nabla^2 + 2F_{ii})^{-1}$. This is achieved first by division by $\boldsymbol{\kappa}$, recalling that $\boldsymbol{\kappa}^{-1} = 1 - V_Z/2c$. The following integral equation is obtained

$$\varphi_i = -2\hat{G}_{\mu_i} \left[-\frac{1}{2} \frac{\nabla \boldsymbol{\kappa}}{\boldsymbol{\kappa}} \cdot \nabla \varphi_i + \left(\frac{V}{\boldsymbol{\kappa}} + \frac{V_Z}{2c^2} F_{ii} \right) \varphi_i - \frac{1}{\boldsymbol{\kappa}} \sum_{j \neq i} F_{ij} \varphi_j \right] \quad (17)$$

When $c \rightarrow \infty$, $\boldsymbol{\kappa} \rightarrow 1$, and $\nabla \boldsymbol{\kappa} \rightarrow 0$, and the nonrelativistic form as in eq 14 is recovered. Equation 17 can therefore be seen as a *level-shifted* version of its nonrelativistic counterpart. Although it cannot be expected that the iterative solution of eq 17 will work for arbitrary shifts, the approach is justified by

recalling that $\kappa \approx 1$ almost everywhere except close to the nuclei. Our tests indicate that it becomes more difficult to converge the above equation when the ZORA contribution becomes larger, either *physically* by going down the periodic table, or *artificially* by letting $c \rightarrow 0$. To overcome the convergence issues for heavier nuclei (fifth row of the periodic table), we have therefore introduced a finite nucleus model, as described in the section **Finite Nucleus Model**. For elements beyond the fifth row, one has to keep in mind that ZORA becomes questionable.⁹¹

An alternative approach to obtain the desired $(\nabla^2 + 2F_{ii})$ term from eq 16 is to add a Laplacian term $(\frac{1}{2}\nabla^2)$ directly, thus avoiding division by κ . Our tests indicate that the strategy presented above works better, despite the additional singularity introduced. The main reason seems to be that the former strategy removes the Laplacian altogether, which is ill-conditioned in the discontinuous MW basis, whereas the latter keeps part of it on the right-hand side.

Implementation within Multiwavelets. An implementation of the ZORA method as outlined above is currently in a development version of the MRChem package⁹² and is expected to appear in the next official release v1.2. MRChem is a numerical quantum chemistry code based on a MW framework, in which all functions and operators are represented on their own fully adaptive multiresolution numerical real-space grid. This allows for efficient all-electron treatment of medium to large molecules (hundreds of atoms) at the self-consistent field (SCF) level of theory (both HF and DFT).

The κ function is computed as a pointwise map of the chosen ZORA potential V_Z through

$$\kappa(r) = \frac{1}{1 - V_Z(r)/2c^2} \quad (18)$$

and similarly for its inverse

$$\kappa^{-1}(r) = 1 - V_Z(r)/2c^2 \quad (19)$$

Both functions are represented on their own adaptive numerical grid and subsequently treated as standard multiplicative potential operators in the SCF procedure of solving eq 17.

Point Nucleus Model. In a MW framework, the singularity of the nuclear potential can lead to numerical problems when V_{nuc} is computed and used. In the nonrelativistic case, the issue is circumvented by replacing the analytic $1/r$ potential with a smoothed approximation⁷⁵

$$u(r) = \frac{\text{erf}(r)}{r} + \frac{1}{3\sqrt{\pi}}(e^{-r^2} + 16e^{-4r^2}) \quad (20)$$

and then parametrized as $u(r/s)/s$, where s is a scalar smoothing parameter. The smoothing parameter depends on the nuclear charge Z and the desired precision ϵ as explained in ref 75.

$$s = \left(\frac{0.00435\epsilon}{Z^5} \right)^{1/3} \quad (21)$$

The above prescription constitutes a *numerical* smoothing of the nucleus to avoid accidental infinity in the representations. It should not be confused with *physical* finite nucleus models,⁹³ which are common in relativistic methods. This numerical smoothing is much sharper than the common finite nucleus

models, and it is meant to yield results that are, within the requested precision, equivalent to using a point charge.

Finite Nucleus Model. In order to overcome the numerical issues faced by the pointwise nuclei, we have introduced the Gaussian nuclear model, as described by Visscher and Dyall.⁹³ Not only does the model overcome the numerical problems of pointwise charges but it is also a sounder physical description of larger nuclei. The nuclear charge is modeled as a Gaussian distribution

$$\rho^G(R) = \rho_0^G e^{-\xi R^2} \quad (22)$$

with the normalization prefactor $\rho_0^G = eZ(\xi/\pi)^{3/2}$ and the parameter ξ related to the root-mean-square radius of the nucleus via the expression $\xi = 3/\langle R^2 \rangle$. To determine $\langle R^2 \rangle$, we apply the empirical formula for ref 94

$$\sqrt{\langle R^2 \rangle} = (0.836A^{1/3} + 0.570)\text{fm} \quad (23)$$

where A is the nuclear mass number and fm is a femtometer length unit (10^{-15} m). We note that a choice of the nuclear mass A is dependent on whether the abundance of isotopes is taken in to account. To avoid the ambiguity, we use the tabulated values of $\sqrt{\langle R^2 \rangle}$ provided in ref 93.

The potential for a Gaussian charge distribution can be represented in an analytic form by Visscher and Dyall.⁹³

$$V^G(R) = -\frac{eZ}{R} \text{erf}(\sqrt{\xi}R) \quad (24)$$

which is the solution to the Poisson equation for the charge density defined by eq 22.

Among the different possibilities presented in ref 93, we have chosen the Gaussian model for the present work because it is simpler than the more realistic Fermi model and because it is implemented in all GTO codes. The goal of the present work is validation of the implementation. It is therefore less relevant which model is used provided that it is consistent throughout the software packages employed.

Methods for Verification. To verify correctness of the ZORA implementation in MRChem, we use two other codes: a numerical atomic solver⁸⁴ and an all-electron full-potential LAPW code EXCITING.¹⁰⁴ In this section, we briefly introduce them and provide details on the implementation of new features needed for making a direct comparison with MRChem.

The atomic solver assumes atoms with spherically symmetric densities, and the mono-electronic wave functions are represented as $\varphi_{nlm}(\mathbf{r}) = u_{nl}(r)Y_{lm}(\hat{r})$, where $u_{nl}(r)$ is a radial function defined on a one-dimensional radial grid. The SCF of atomic solver uses an approach similar to that of the MW framework described above. It reduces the nonrelativistic Kohn–Sham equation to the integral form as follows

$$\varphi_{nlm}(\mathbf{r}) = -2\hat{G}_{e_n}(\hat{V}\varphi_{nlm}(\mathbf{r})) \quad (25)$$

This equation matches eq 14, with vanishing off-diagonal terms of the Fock matrix because the canonical representation is used. The radial solver originally supported only non-relativistic calculations: to implement ZORA, we used eq 17 in the canonical form, i.e., replacing F_{ii} with e_{n_i} and setting $F_{ij} = 0$ for $i \neq j$. This approach avoids the evaluation of second derivatives and the corresponding numerical noise, which accumulates during the self-consistency iterations.

To consider systems beyond atoms, we use EXCITING code. In a nutshell, the code relies on partitioning the unit cell into nonoverlapping atomic spheres and the interstitial region. In the atomic spheres, the wave functions are expressed in terms of atomic-like orbitals that are updated during the self-consistency cycle. In the interstitial region, one represents the wave functions with plane waves using the smoothness of the Kohn–Sham potential. Based on such an approach, we use two types of basis functions: (i) augmented plane-waves (APWs)^{95,96} and (ii) local-orbitals (LOs).⁹⁷ Each APW combines a plane wave in the interstitial region with atomic-like orbitals in the spheres, whereas each LO is a linear combination of two atomic-like orbitals in one particular sphere and strictly zero everywhere else. As shown in ref 86, it is possible to obtain a systematic convergence of the total energies simply by increasing the number of APWs and LOs.

The ZORA Hamiltonian is already available with the point nucleus model in the released version of EXCITING code in 2021, whereas the Gaussian nuclear model charge feature described in the **Finite Nucleus Model** section is implemented in the present work. Aside from adopting the electrostatic potential due to a Gaussian charge density, we also revise the evaluation of the following integral

$$E_M = \int \rho_n(\mathbf{r}) V_C(\mathbf{r}) d\mathbf{r} \quad (26)$$

If a nucleus at the site \mathbf{R}_0 is defined as a point charge, then its density distribution is $\rho_n(\mathbf{R}) = Z\delta(\mathbf{R} - \mathbf{R}_0)$ leading to $E_M = ZV_C(\mathbf{R}_0)$. For smeared nuclei with $\rho_n(\mathbf{R}) = \rho^G(\mathbf{R})$ (see eq 24), eq 26 is evaluated as an integral on a radial grid.

COMPUTATIONAL DETAILS

All calculations were performed with the PBE functional⁹⁸ using the XCFun⁹⁹ library, the LibXC¹⁰⁰ library, and the native implementation in MRChem, the atomic solver, and EXCITING, respectively. Being aware of slight inconsistencies in the PBE parameters employed in these libraries, we set $\beta = 0.06672455060314922$ and $\mu = 0.066725\frac{\pi^2}{3}$ as defined in XCFun and use these values in calculations with all three codes.

For all MW calculations, a development version of MRChem has been employed. Interpolating polynomials of ninth order were used with a simulation box size of $128 a_0$. The numerical precision thresholds used are summarized in Table 1.

Table 1. MRChem Precision Parameters

parameter	value	explanation
world_prec	1.0×10^{-6}	overall numerical precision
energy_thrs	1.0×10^{-6}	convergence threshold total energy
orbital_thrs	1.0×10^{-4}	convergence threshold maximum orbital residual

All EXCITING calculations were performed using a large cubic unit cell with a side length of $25 a_0$ for atoms and $30 a_0$ for molecules. The electrostatic interaction of the periodic images was eliminated by introducing the truncation of the Coulombic interaction following the approach explained in ref 81. Aside from acquiring the isolated limit, this adjustment allows us to obtain ZORA energies consistent with both MRChem and the atomic solver. The unmodified electrostatic potential for periodic densities is defined uniquely except for

an additive constant which introduces an ambiguity in the ZORA energies.^{34,37,54–56} The truncation of the Coulombic interaction makes the potential unique and thus removes this ambiguity completely.^{34,37,54–56} The canonical orbitals were expressed in terms of local orbitals and LAPWs with the cutoff $R_{MT}K_{max}$ sufficient to ensure a few μ Ha precision. The specific settings in the case of each atom and molecule are stored in Table 2, and a set of input and output data files are available in the repository.

Table 2. LAPW Parameters and Structural Data Used in the Calculations of the Considered Molecules and Atoms^a

material	$R_{MT}^{min}G_{max}$	$R_{MT} [a_0]$	bond length [Å]
CaO	11	1.63/1.40	1.8221 ¹⁰¹
CuH	10	1.40/1.00	1.4626 ¹⁰¹
SrO	13	1.63/1.40	1.9050 ¹⁰²
Cu ₂	14	1.8	2.2197 ¹⁰¹
AgH	11	1.68/1.20	1.6180 ¹⁰¹
I ₂	14	1.6	2.6630 ¹⁰²
He	11	2	
Ne	12	2	
Ar	13	2	
Kr	13	2	
Xe	15	2	

^a $R_{MT}^{min}G_{max}$ is the product of the smallest muffin-tin radius and the largest reciprocal lattice vector.

The calculations with the atomic solver employed a fifth order polynomial on a radial grid with the innermost and outermost points $r_{min} = 10^{-8} a_0$ and $r_{max} = 35 a_0$. The number of radial points was set to 5000, which is fully sufficient to guarantee sub- μ Ha precision.

The total energies of diatomic molecules were calculated in MRChem and EXCITING without geometry relaxation using the internuclear distances given in Table 2. The bond lengths in Table 2 are taken from the NIST¹⁰¹ database with the exception of SrO and I₂, which are private communication from not published work.¹⁰²

RESULTS AND DISCUSSION

Relative Contributions of the ZORA Terms. The current MRChem implementation is able to include all local contributions of the potential into \hat{V}_Z . It is therefore possible to measure their relative weights for a given atom and how their contribution changes with nuclear charge. We have performed a series of calculations for the noble gases from helium to xenon, with all seven possibilities: one contribution only, two contributions, and all three. The results are summarized in Figure 1. The nuclear potential \hat{V}_{nuc} is the largest contribution, as expected. It is followed by the Coulombic term \hat{J} , and the exchange and correlation potential \hat{V}_{xc} is the smallest one. Moreover, the relativistic correction increases roughly with the fourth power of the nuclear charge as expected,⁵ and the nuclear term becomes progressively more dominant for heavier atoms.

Validation. To validate the ZORA implementation in MRChem, we perform total energy calculations of noble gas atoms and a small set of molecules broadly covering the first five rows of the periodic table (H–Xe). In the case of the atoms, we apply three different types of calculations: the atomic solver, LAPWs, and MWs. For the diatomic molecules,

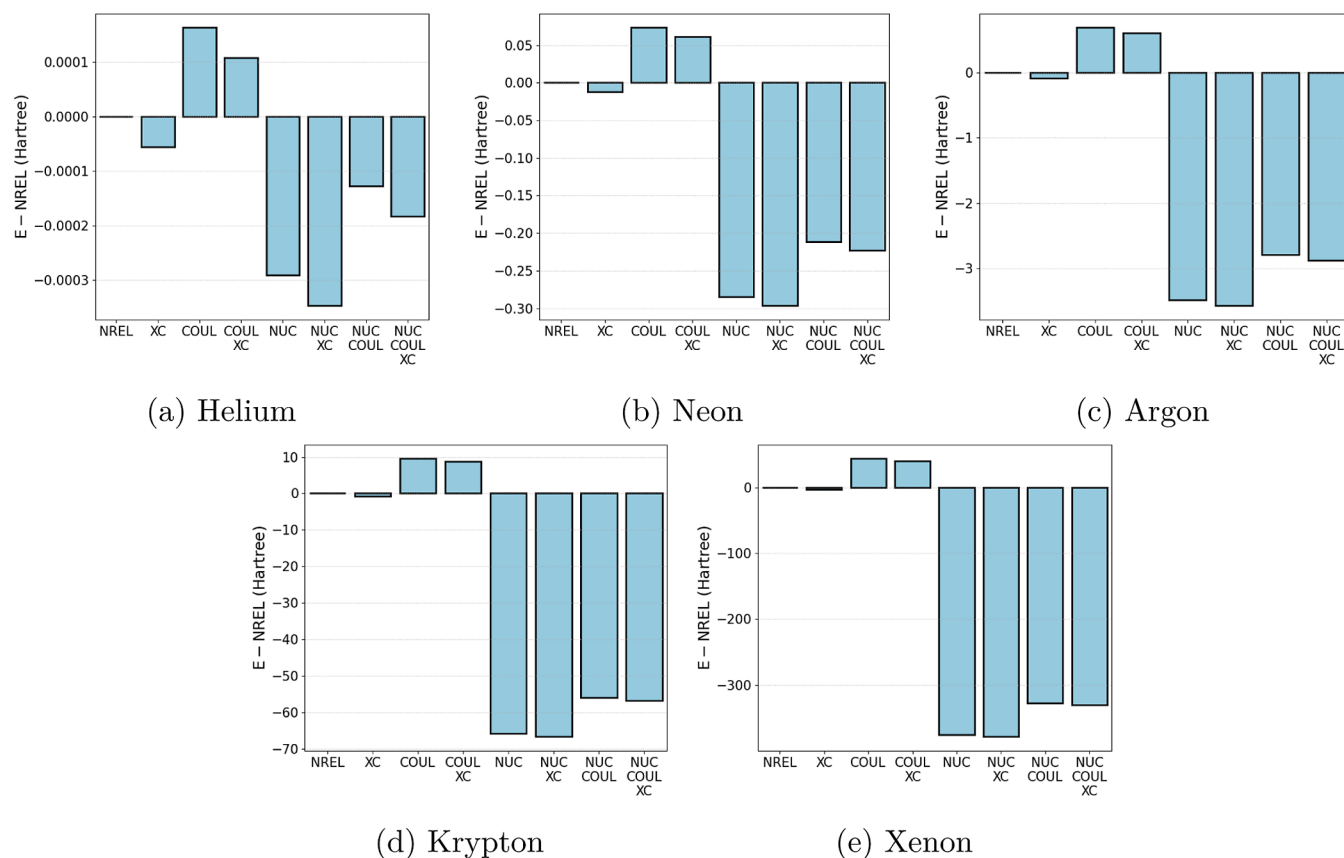


Figure 1. Relative contributions to the total energy (in Hartree) for all seven possible ZORA operators, compared to the nonrelativistic total energy, for the noble gases He–Xe. Note that the scales on the y -axes are different for each subplot.

Table 3. Non-Relativistic Total Energies (Given in Hartrees) Obtained Using Three Different Codes and Their Relative Differences^a

species	atomic solver	EXCITING	MRChem	relative difference	
				MRChem vs atomic solver	MRChem vs EXCITING
He	-2.892935	-2.892935	-2.892935	0.0	0.0
Ne	-128.866434	-128.866433	-128.866433	-3.1×10^{-10}	0.0
Ar	-527.346141	-527.346138	-527.346140	-1.7×10^{-9}	3.7×10^{-9}
Kr	-2753.416138	-2753.416137	-2753.416138	0.0	2.1×10^{-10}
Xe	-7234.233259	-7234.233259	-7234.233259	0.0	0.0
CaO		-752.562050	-752.562058		1.0×10^{-8}
CuH		-1640.901717	-1640.901727		6.1×10^{-9}
SrO		-3208.161712	-3208.161722		3.1×10^{-9}
Cu ₂		-3280.675840	-3280.675846		1.7×10^{-9}
AgH		-5200.162245	-5200.162259		2.8×10^{-9}
I ₂		-13840.158307	-13840.158328		1.5×10^{-9}

^aIn all cases, the point-like nucleus model is used.

we compare the results obtained using MW and LAPW by means of MRChem and EXCITING, respectively.

To assess the potential agreement which can be achieved between the various codes, we have performed nonrelativistic calculations with all three codes, using point-charge nuclei (numerically smoothed as described in [Point Nucleus Model](#) section in the case of MWs). The results are summarized in [Table 3](#). We have then obtained a root-mean-square deviation (rmsd) of the relative error between MWs and LAPWs equal to 5.21×10^{-8} and between MWs and atomic solver equal to 8.07×10^{-10} . We concluded that the three methods are in very good agreement, setting the mark for what can be expected in

the ZORA domain, using a Gaussian nuclear charge distribution as in [eq 24](#). The ZORA results are summarized in [Table 4](#), showing a rmsd for the relative errors between MWs and LAPWs equal to 3.95×10^{-8} , and between MWs and atomic solver equal to 6.93×10^{-9} , in line with what has been here previously observed for the nonrelativistic regime, thus confirming the validity of the implementation. In terms of absolute errors, we find that most discrepancies in the total energies are within 10 microHartrees, and only in two cases (AgH and I₂) the differences are larger, yet they do not exceed 21 microHartrees. This level of agreement is much better than the so-called *chemical accuracy threshold* (1 kcal/mol).

Table 4. Scalar-Relativistic ZORA Total Energies (Given in Hartrees) Obtained Using Three Different Codes and Their Relative Differences^a

species	atomic solver	EXCITING	MRChem	relative difference	
				MRChem vs atomic solver	MRChem vs EXCITING
He	-2.893119	-2.893119	-2.893119	0.0	0.0
Ne	-129.089335	-129.089335	-129.089335	0.0	0.0
Ar	-530.224039	-530.224036	-530.224038	-2.6×10^{-9}	2.9×10^{-9}
Kr	-2810.049764	-2810.049762	-2810.049755	-3.1×10^{-9}	-2.4×10^{-9}
Xe	-7564.596665	-7564.596662	-7564.596554	-1.5×10^{-8}	-1.4×10^{-8}
CaO		-757.195492	-757.195491		-1.7×10^{-9}
CuH		-1663.238493	-1663.238504		6.7×10^{-9}
SrO		-3279.826587	-3279.826596		2.7×10^{-9}
Cu ₂		-3325.342851	-3325.342843		-2.4×10^{-9}
AgH		-5380.042000	-5380.042003		5.5×10^{-10}
I ₂		-14448.747787	-14448.747761		-1.8×10^{-9}

^aIn all cases, the smeared nucleus model is used.

Finally, we find that introducing the relativistic corrections in the MW formalism, as expressed in eq 17, leads to a minor increase in MRChem runtimes. Therefore, we anticipate that the analysis of the performance (runtimes and parallel scaling) given in ref 103 remains valid in the ZORA case.

CONCLUSIONS

We have formulated the ZORA method in a form compatible with MWs and implemented it in the MRChem program. The validity and precision of the implementation have been tested against a radial, numerical atomic code and a plane wave code, showing excellent agreement. The current study was performed with the specific idea to validate method and theory with a small benchmark: atoms and diatomics, covering broadly the periodic table up to and including fifth-row elements. The model is also capable of dealing with all parts of the electronic potential self-consistently, with the exception of the HF exchange, which is the only one that cannot be expressed in the closed form.

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